DEPARTMENT OF CIVIL ENGINEERING UNIVERSITY OF SOUTHERN CALIFORNIA

PRELIMINARY EMPIRICAL MODEL FOR SCALING FOURIER AMPLITUDE

SPECTRA OF STRONG GROUND ACCELERATION IN TERMS OF EARTHQUAKE MAGNITUDE

SOURCE TO STATION DISTANCE, SITE INTENSITY AND RECORDING SITE CONDITIONS

by

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ABSTRACT

In this paper, we present the current improvements in empirical scaling of Fourier spectrum amplitudes of strong earthquake accelerations by introducing the frequency dependent attenuation function which has been developed (Trifunac and Lee, 1985) from the same data base. This function replaces the Richter's empirical attenuation function which we previously used together with a linear term in R, the epicentral distance. By using the new attenuation function, the scaling model has the additional flexibility for estimating the Fourier spectral amplitudes from earthquakes of given source dimensions and focal depths.

INTRODUCTION

The idea to scale Fourier amplitude spectra of strong earthquake ground motion directly in terms of earthquake magnitude or Modified Mercalli Intensity at a site is not new and has been considered by a number of investigators during the past ten years. In fact, the basic ideas and equations employed here are essentially the same as those we presented in 1976 (Trifunac, 1976b). With recent significant increase in the number of uniformly processes strong motion accelerograms however, it has been possible to detect frequency and source size dependent trends in the spectrum amplitude attenuation with distance. Because such refinement of attenuation laws should lead to smaller scatter of observed Fourier spectrum amplitudes about the empirical scaling models and thus to more reliable estimates of the amplitudes of future strong ground motion, the aim of this paper is to present the second generation of "preliminary empirical scaling models..." (Trifunac, 1976b) by including this new description of amplitude attenuation with distance.

In some respects the analyses such as this one will continue to be of preliminary nature so long as the new data will contribute new and significant changes in the functional forms of the regression equations. With the overall empirical scaling models slowly "converging" towards an accurate and complete representation, it is hoped that the present effort will contribute one such improvement.

In this analysis we continue to employ the "published" magnitude scale to describe the size of earthquakes in our data base (Trifunac and Lee 1985). With increasing number of well studied earthquakes for which strong motion data are available, it should be possible, in the near future, to develop similar empirical scaling relations, but in

terms of more "physical" earthquake source parameters, for example, seismic moment and stress drop. Such scaling parameters are expected to decrease the overall fluctuations of the recorded amplitudes about the average empirical estimates. However, one of the principal uses of the scaling models presented here will continue to be for the calculation of Uniform Risk Spectra (Anderson and Trifunac, 1978). Such probabilistic estimates of strong ground motion are still, in many parts of the world, based on old seismicity records where in some cases even the estimates of earthquake magnitude have to be derived from old and often incomplete data on reported intensities of shaking. While the empirical relations between the seismic moment, magnitude and the reported intensities are available, it is not clear, at present, how much could be gained by converting all scaling relationships from magnitude to seismic moment, for example.

In defining the "distance" between the earthquake source and the recording stations, in this work we consider approximately the effects of source depth and source size, but we continue to employ the epicentral distance to define the principal horizontal distance component. One could consider, instead the closest distance to a fault or distance perpendicular to the fault projection on ground surface. Such distance definitions would imply that some information on the distribution of energy release along the fault is available. Since this is available only for a small subset of earthquakes contributing to the data base considered here, we chose to continue with the simplified distance definition in terms of the epicentral distance.

PART I: SCALING OF FOURIER SPECTRA IN TERMS OF M, R, H, S, h and v
I.1 PREVIOUS ANALYSIS

During the regression analyses of earthquake strong-motion parameters in the 1970's Trifunac (1976b) suggested that the Fourier amplitude spectra (FS) of strong motion acceleration at a selected set of discrete periods, T, can be scaled in terms of the definition of the earthquake magnitude scale and a "correction" function in the following form:

 $\log_{10}[FS(T)_{p}] = M + \log_{10}A_{o}(R) - \log_{10}\{FS_{o}(T,M,p,s,v,R)\},$ where M is the local earthquake magnitude, M_L ; $log_{10}A_o(R)$ represents the amplitude attenuation function (Richter, 1958) versus distance (Table I.1.1). The term $log_{10}{FS_0(T,M,p,s,v,R)}$ represents a "correction" function which incorporates the effects of: (1) distribution of observations with respect to the assumed empirical model, as represented by the confidence level p selected for the approximate bound of spectral amplitudes $FS(T)_{p}$, (2) geologic site conditions, s, (s=0 for alluvium, s=2 for basement rock, s=1 for intermediate sites), (3) horizontal versus vertical ground motion differences, (v=0 for horizontal and v=1 for vertical), and (4) the frequency dependent attenuation effects of amplitudes versus distances, R. The term $\log_{10}A_{o}(R)$ was empirically determined for Southern California (Richter, 1958) and is representative of wave frequencies centered near the middle of the frequency band for the data obtained from digitization and data processing of strong motion accelerograms (0.1 Hz to 25 Hz). The term $log_{10}{FS_0(T,M,p,s,v,R)}$ was then determined by regression analysis. The same empirical model was also used for scaling of pseudo relative velocity spectra, PSV (Trifunac and Anderson, 1978a) and relative velocity spectra, SV (Trifunac and Anderson, 1978b).

Table I.1.1 $\log_{10}A_o(R)$ vs epicentral distance R

R(km)	-log _{lo} A _o (R)	R(km)	-log _{lO} A _o (R)	R(km)	-log _{lO} A _o (R)
1	1.400	140	3.230	370	4.336
5	1.500	150	3.279	380	4.376
10	1.605	160	3.328	390	4.414
15	1.716	170	3.378	400	4.451
20	1.833	180	3.429	410	4.485
25	1.955	190	3.480	420	4.518
30	2.078	200	3.530	430	4.549
35	2.199	210	3.581	440	4.579
40	2.314	220	3.631	450	4.607
45	2.421	230	3.680	460	4.634
50	2.517	240	3.729	470	4.660
55	2.603	250	3.779	480	4.685
60	2.679	260	3.827	490	4.709
65	2.746	270	3.877	500	4.732
70	2.805	280	3.926	510	4.755
80	2.920	290	3.975	520	4.776
85	2.958	300	4.024	530	4.797
90	2.989	310	4.072	540	4.817
95	3.020	320	4.119	550	4.835
100	3.044	330	4.164	560	4.853
110	3.089	340	4.209	570	4.869
120	3.135	350	4.253	580	4.885
130	3.182	360	4.295	590	4.900

Trifunac and Lee (1978) refined the above analyses by introducing a measure of the depth of sedimentary deposits beneath the recording station, h, as a site characteristic to replace the scaling parameter s mentioned above. The new scaling equation then became (equation (1) of Trifunac and Lee (1978)):

$$log_{10}[FS(T)] = M + log_{10}A_o(R) - b(T)M - c(T) - d(T)h$$

- $e(T)v - f(T)M^2 - g(T)R$, (I.1.2)

with all the parameters defined as above. The functions b(T), c(T), ..., and g(T) are estimated by regression analysis at 91 periods T between 0.04 sec and 15 sec.

Note that in the regression equation (I.1.2) the second and higher order terms of h, R, and the third and higher order terms of M are neglected. It was pointed out then that there is really no physical basis to assume that $\log_{10}FS(T)$ should be just a linear function of the depth of sediments. Before choosing the final form for equation (I.1.2) then, studies were carried out to find whether there is any significant dependence of spectral amplitudes on h^2 , h^3 , ... etc. It was found that with the database available then, the least-squares coefficients associated with these higher order terms of h are indistinguishable from zero at the 95% confidence level. It should be noted, also, that there is no physical justification for the chosen parabolic dependence on M. This choice is motivated by the simplicity of its mathematical form and the apparent trend of data indicated in earlier analyses (Trifunac and Brady, 1975).

Note that if the $\log_{10}A_o(R)$ in (I.1.2) were to represent the geometric spreading, the term g(T)R would model the equivalent anelastic attenuation. However, $\log_{10}A_o(R)$ was derived empirically from data on actual peak amplitudes in Southern California, and thus represents an average combination of geometric spreading and anelastic attenuation. The term g(T)R then only represents a correction to the average attenuation given by $\log_{10}A_o(R)$.

I.2 THE NEW DATABASE

The above regression analysis was carried out for 186 free-field records corresponding to a total of 558 components of data from 57 earthquakes starting with the Long Beach earthquake in 1933 and ending with the San Fernando earthquake in 1971. Through the years new earthquake acceleration data have been added to the original database. The list of 57 earthquakes has now grown to 104, most of which occurred in the regions of northern and southern California. Table I.2.1 is the list of earthquakes now used in our database. Each line contains information on the date and time of the earthquake, latitude and longitude of the epicenter, focal depth, local earthquake magnitude and maximum intensity, if available, and the name of the earthquake.

The original list of 186 free-field records corresponding to 57 earthquakes has now grown to 438 free-field records from these 104 earthquakes. With 3 components available for each record, this amounts to a total of 1314 acceleration components, of which there are 876 horizontal and 438 vertical components.

```
EQ MON/DAY/YR TIME
                      LATITUDE LONGTITUDE (KM)
                                                   MAX
                                                         NAME
               CODE
                        DEG, MIN & SEC
                                         DEPTH MAG MMI
   3 10 1933 1754PST 33 37 00 -117 58 00 16.0 6.3 9 LONG BEACH, CALIF
 2 10  2 1933 0110PST 33 47 00 -118 08 00 16.0 5.4 6 SOUTHERN CALIF
       6 1934 1449PST 41 42 00 -124 36 00
                                                   5 EUREKA, CALIF
 4 12 30 1934 0552PST 32 15 00 -115 30 00 16.0 6.5 9 LOWER CALIF
 5 10 31 1935 1138MST 46 37 00 -111 58 00
                                               6.0 8 HELENA, MT
 6 10 31 1935 1218MST 46 37 00 -111 58 00
                                                   3 HELENA, MT
 7 11 21 1935 2058MST 46 36 00 -112 00 00
                                                   6 HELENA, MT
 8 11 28 1935 0742MST 46 37 00 -111 58 00
                                                   6 HELENA, MT
      6 1937 2042PST 40 30 00 -125 15 00
                                                   5 HUMBOLDT BAY, CAL
    4 12 1938 0825PST 32 53 00 -115 35 00 16.0 3.0 IMPERIAL VALLEY, CA
       5 1938 1842PST 32 54 00 -115 13 00 16.0 5.0 IMPERIAL VALLEY, CA
11
      6 1938 0435PST 32 15 00 -115 10 00 16.0 4.0 IMPERIAL VALLEY, CA
12
    9 11 1938 2210PST 40 18 00 -124 48 00
                                               5.5 6 NW CALIF
    5 18 1940 2037PST 32 44 00 -115 30 00 16.0 6.710 IMPERIAL VALLEY, CA
      9 1941 0145PST 40 42 00 -125 24 00
                                              6.4
                                                     NW CALIF
   6 30 1941 2351PST 34 22 00 -119 35 00 16.0 5.9 8 SANTA BARBARA, CAL
      3 1941 0813PST 40 36 00 -124 36 00
                                               6.4 7 NORTHERN CALIF
18 11 14 1941 0042PST 33 47 00 -118 15 00 16.0 5.4 8 TORRANCE-GARDENA CA
19 10 21 1942 0822PST 32 58 00 -116 00 00 16.0 6.5 7 BORREGO VALLEY, CAL
20 3 9 1949 0429PST 37 06 00 -121 18 00
                                              5.3 7 NORTHERN CALIF
    4 13 1949 1156PST 47 06 00 -122 42 00
21
                                              7.1 8 WESTERN WASH
   1 23 1951 2317PST 32 59 00 -115 44 00 16.0 5.6 7 IMPERIAL VALLEY, CA
23 10 7 1951 2011PST 40 17 00 -124 48 00
                                               5.8 7 NW CALIF
    7 21 1952 0453PDT 35 00 00 -119 01 00 16.0 7.711 KERN COUNTY, CALIF
25
    7 23 1952
                      35 17 00 -118 39 00
                                                     KERN CNTY, CAL
   9 22 1952 0441PDT 40 12 00 -124 25 00
                                               5.5 7 NORTHERN CALIF
27 11 21 1952 2346PST 35 50 00 -121 10 00
                                              6.0 7 SOUTHERN CALIF
   6 13 1953 2017PST 32 57 00 -115 43 00 16.0 5.5 7 IMPERIAL VALLEY, CA
   1 12 1954 1534PST 35 00 00 -119 01 00 16.0 5.9 8 WHEELER RIDGE, CALI
30 4 25 1954 1233PST 36 48 00 -121 48 00
                                               5.3 7 CENTRAL CALIF
31 11 12 1954 0427PST 31 30 00 -116 00 00 16.0 6.3 5 LOWER CALIF
32 12 21 1954 1156PST 40 47 00 -123 52 00
                                              6.5 7 EUREKA, CALIF
     4 1955 1801PST 37 22 00 -121 47 00
                                          5.8 7 SAN JOSE, CALIF
34 12 16 1955 2117PST 33 00 00 -115 30 00 16.0 4.3
                                                     IMPERIAL COUNTY, CA
35 12 16 1955 2142PST 33 00 00 -115 30 00 16.0 3.9
                                                     IMPERIAL COUNTY, CA
36 12 16 1955 2207PST 33 00 00 -115 30 00 16.0 5.4 7 IMPERIAL COUNTY
      9 1956 0633PST 31 42 00 -115 54 00 16.0 6.8
                                                   EL ALAMO, BAJA CAL
      9 1956 0725PST 31 42 00 -115 54 00
38
                                              6.4
                                                     EL ALAMO, BAJA CAL
   3 18 1957 1056PST 34 07 06 -119 13 12 13.8 4.7 6 SOUTHERN CALIF
   3 22 1957 1048PST 37 40 00 -122 28 00
                                              3.8 5 SAN FRANCISCO CA
41
   3 22 1957 1144PST 37 40 00 -122 29 00
                                              5.3 7 SAN FRANCISCO, CAL
42
   3 22 1957 1515PST 37 39 00 -122 27 00
                                              4.4 5 SAN FRANCISCO CA
   3 22 1957 1627PST 37 39 00 -122 29 00
43
                                              4.0 5 SAN FRANCISCO CA
   1 19 1960 1926PST 36 47 00 -121 26 00
                                              5.0 6 CENTRAL CALIF
45
      5 1960 1718PST 40 49 00 -124 53 00
                                              5.7 6 NORTHERN CALIF
      8 1961 2323PST 36 30 00 -121 18 00 11.0 5.7 7 HOLLISTER, CALIF
46
47
      4 1962 0917PST 40 58 00 -124 12 00
                                              5.0 6 NORTHERN CALIF
48
   4 29 1965 0729PST 47 24 00 -122 18 00
                                              6.5 8 PUGET SOUND, WASH
   7 15 1965 2346PST 34 29 06 -118 31 18 15.1 4.0 6 SOUTHERN CALIF
49
50
   6 27 1966 2026PST 35 57 18 -120 29 54 6.0 5.6 7 PARKFIELD. CALIF
      7 1966 0936PST 31 48 00 -114 30 00 16.0 6.3 6 GULF OF CALIF
51
   9 12 1966 0841PST 39 24 00 -120 06 00 6.3 7 NORTHERN CALIF
53 12 10 1967 0407PST 40 30 00 -124 36 00
                                          5.8 6 NORTHERN CALIF
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54 12 18 1967 0925PST 37 00 36 -121 47 18
                                            5.2 6 NORTHERN CALIF
        8 1968 1830PST 33 11 24 -116 07 42 11.1 6.4 7 BORREGO MTN, CALIF
     9 12 1970 0630PST 34 16 12 -117 32 24 8.0 5.4 7 LYTLE CREEK, CALIF
        9 1971 0600PST 34 24 42 -118 24 00 13.0 6.411 SAN FERNANDO, CALIF
 58 10 15 1979 1417PST 32 37 59 -115 19 59 12.0 6.6
                                                      IMPERIAL VALLEY, CA
        6 1979 0805PST 37 06 43 -121 31 59 9.6 5.9
                                                       COYOTE LAKE, CALIF
     8 13 1978 2254GMT 34 21 04 -119 42 00 12.5 5.5
                                                       SANTA BARBARA, CAL
 61
     1 24 1980 1100PST 37 49 37 -121 47 13
                                            5.9 5.9
                                                      MT. DIABLO, LIVERMO
     1 26 1980 1833PST 37 45 00 -121 42 47
                                                      MT. DIABLO, LIVERMO
                                            7.3 5.2
 63 08 02 1975 2022GMT 39 26 58 -121 28 25
                                            4.1 5.2
                                                      OROVILLE AFTERSHOCK
 64 08 02 1975 2059GMT 39 26 00 -121 28 31
                                            5.1 5.2
                                                      OROVILLE AFTERSHOCK
 65 08 03 1975 0103GMT 39 29 19 -121 30 59
                                            8.8 4.6
                                                      OROVILLE AFTERSHOCK
 66 08 03 1975 0247GMT 39 28 52 -121 30 21
                                            7.4 4.1
                                                      OROVILLE AFTERSHOCK
 67 08 05 1975 0228GMT 39 24 18 -121 29 43
                                            6.2 3.2
                                                      OROVILLE AFTERSHOCK
 68 08 06 1975 0350GMT 39 29 46 -121 31 49
                                            9.2 4.7
                                                      OROVILLE AFTERSHOCK
 69 08 06 1975 1641GMT 39 29 31 -121 31 45
                                            9.7 3.9
                                                      OROVILLE AFTERSHOCK
 70 08 08 1975 0700GMT 39 29 50 -121 30 41
                                            7.7 4.8
                                                      OROVILLE AFTERSHOCK
 71 08 11 1975 0611GMT 39 27 29 -121 28 59
                                            3.1 4.4
                                                      OROVILLE AFTERSHOCK
 72 08 11 1975 1559GMT 39 30 20 -121 31 35
                                            9.8 3.8
                                                      OROVILLE AFTERSHOCK
 73 08 16 1975 0548GMT 39 28 12 -121 31 42
                                                      OROVILLE AFTERSHOCK
                                           8.5 4.1
 74 08 16 1975 1223GMT 39 29 52 -121 30 16
                                           7.1 3.1
                                                      OROVILLE AFTERSHOCK
 75 09 27 1975 2234GMT 39 31 12 -121 31 56 10.4 4.6
                                                      OROVILLE AFTERSHOCK
 76 11 28 1974 2301GMT 36 54 0 -121 30 0
                                           9.0 0. 6 HOLLISTER, CAL
 77
     1 11 1975 1737PST 40 13 12 -124 15 36
                                           2.0 4.7 6 NORTHERN CAL
 78
        6 1975 1835PST 40 16 48 -124 40 12
                                           0. 4.0
                                                      NORTHERN CAL
 79
        7 1975 0846GMT 40 34 12 -124 08 24 21.0 5.7 7 NORTHERN CAL
 80
     3
        8 1971 1508PST 35 40
                             0 -118 24 12 6.0 4.7 5 CENTRAL CAL
        2 1971 0608GMT 51 24
 81
                             0 -177 12 0 43.0 7.1 6 ANDREANOF, ALASKA
    9 12 1971 1132PST 41 17 54 -123 40 24 20.0 4.6 5 NORTHERN CAL
 82
    7 30 1972 2145GMT 56 49 12 -135 40 48 25.0 7.1 7 SOUTHEAST ALASKA
 83
 84
        4 1972 1804GMT 36 38 13 -121 17 13
                                           2.0 4.8 6 CENTRAL CAL
    5 26 1980 1857GMT 37 32 37 -118 51 41
 85
                                           2.8 4.9
                                                      MAMMOTH AFTERSHOCK
    5 27 1980 1450GMT 37 27 49 -118 49 24
 86
                                            2.4 6.3
                                                      MAMMOTH AFTERSHOCK
     5 27 1980 1901GMT 37 36 15 -118 46 11
 87
                                            3.8 5.0
                                                      MAMMOTH AFTERSHOCK
 88
      28 1980 0516GMT 37 34 49 -118 53 09
                                            3.3 4.8
                                                      MAMMOTH AFTERSHOCK
 89
    5 31 1980 1516GMT 37 32 22 -118 54 22
                                           8.2 5.1
                                                      MAMMOTH AFTERSHOCK
    6 11 1980 0441GMT 37 30 24 -119 02 34 14.1 5.0
 90
                                                      MAMMOTH AFTERSHOCK
 91
     6 28 1980 0058GMT 37 33 23 -118 51 45
                                            5.1 4.1
                                                      MAMMOTH AFTERSHOCK
 92 10 16 1979 1616PDT 33
                           4 29 -115 33 16
                                            5.0 4.9
                                                      IMPERIAL VALLEY AFT
 93 10 16 1979 1445PDT 33
                           2 44 -115 29 24
                                            3.9 4.6
                                                      IMPERIAL VALLEY AFT
 94 10 16 1979 1114PDT 32 58 19 -115 36 22
                                            4.7 4.2
                                                      IMPERIAL VALLEY AFT
 95 10 15 1979 2319GMT 32 46 0 -115 26 29
                                            9.5 5.0
                                                      IMPERIAL VALLEY AFT
 96
    4 26 1981 1209GMT 33
                          7 48 -115 39 0
                                           8.0 5.6
                                                      WESTMORELAND, CAL
 97
     1 24 1980 1900GMT 37 50 24 -121 48
                                            5.9 5.9
                                        0
                                                      LIVERMORE, CAL
 98
     1 26 1980 0233GMT 37 45 36 -121 42 0
                                           7.3 5.2
                                                      LIVERMORE, CAL
99
      25 1980 0934PDT 37 36 32 -118 50 49
                                            9.0 6.1
                                                      MAMMOTH AFTERSHOCK
     5 25 1980 0949PDT 37 37 41 -118 55 37
100
                                            14. 6.0
                                                      MAMMOTH AFTERSHOCK
     5 25 1980 1245PDT 37 33 40 -118 49 52
101
                                            16. 6.1
                                                      MAMMOTH AFTERSHOCK
102
     5 25 1980 1336PDT 37 37 30 -118 51 32
                                             2. 5.7
                                                      MAMMOTH AFTERSHOCK
     5 26 1980 1158PDT 37 32 35 -118 53 17
103
                                            5. 5.7
                                                      MAMMOTH AFTERSHOCK
     5 27 1980 0751PDT 37 30 22 -118 49 34
104
                                           14. 6.2
                                                      MAMMOTH AFTERSHOCK
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I.3 THE NEW ATTENUATION FUNCTION

The advantage in using the attenuation function, $\log_{10}A_o(R)$ (Richter, 1958), in the previous regression analyses has been that it contains information on the average properties of wave propagation through the crust in southern California, where virtually all strong motion data have been recorded up to and during the 70's. The disadvantages and limitations have been that its shape does not depend on the magnitude, source dimension and focal depth of an earthquake, on the geological environment of the recording station, or on the amplitudes of the recorded motions. That $\log_{10}A_o(R)$ or its analog should depend on the geometric size of the fault has been discussed in some detail previously (Trifunac, 1976b). Up to the 1970's only a few of the 186 records had epicentral distances less than 10 km and the empirical derivation of different shapes of $\log_{10}A_o(R)$, or its equivalent, to reflect different magnitudes or source dimensions then was not feasible.

With the new database now available, Trifunac and Lee (1985) have developed an iteration procedure for determining a new frequency dependent attenuation function, a complete description of which is given in the above reference. A brief summary and description of this new attenuation function is given here.

To take into account that the attenuation function should depend on the epicentral distance, R, on the focal depth, H, of the earthquake, and the "size" of the fault, S, a parameter, denoted by Δ , is introduced to replace the epicentral distance, R, and is defined as follows:

$$\Delta = S / \left(\ln \left(\frac{S^2 + R^2 + H^2}{S_0^2 + R^2 + H^2} \right) \right)^{\frac{1}{2}}$$
 (1.3.1)

 Δ can be thought of as a "representative distance" from the earthquake source of size S, at depth H and at distance R from the recording site. S_0 is the coherence radius of the source. The definition of Δ used here in equation I.3.1 is identical to that used in Model III (equation 4.8) presented in Trifunac and Lee (1985). It has been proposed by Gusev (1983) in his descriptive statistical model of earthquake source radiation for the description of short-period strong ground motion. The coherence radius S_0 is taken to be a half of the wavelength, λ , for radiation of frequency f (or period T), namely, coherence radius, $S_0 = \lambda/2 = C_{\rm S}/2{\rm f} = C_{\rm S}T/2$, where $C_{\rm S}$ is the velocity of the radiation (in this work $C_{\rm S}$ is taken to be 1 km/sec). Since the fault size, S, of the earthquake is not available for most of the earthquakes used in the data base, an empirical formula for the size, as a function of magnitude, epicentral distance R and period of the spectral amplitudes has been introduced as follows,

$$S = S(M,T)$$
 (1.3.2)

where S(M,T) is the size of the fault "felt" at the period T, and is assumed to be a linear function of magnitude, M, so that for

$$M = 3$$
 $S(M,T) = 0.2 \text{ km}$ $M = 6.5$ $S(M,T) = S_{6.5}(T) \text{ km}$ (I.3.3)

and $S_{6.5}(T)$ is an empirically determined function. From (I.3.3), S(M,T) takes the form

$$S(M,T) = 0.2 + \frac{(M-3)}{3.5} (S_{6.5}(T) - .2)$$
 (I.3.4)

This definition of fault size "felt or "experienced" at the site is independent of how close the site is to the epicenter of the earthquake.

The new frequency dependent attenuation function then takes the form (Trifunac and Lee, 1985):

$$\mathcal{A}tt(\Delta, M, T) = \begin{cases} \mathcal{A}_{o}(T)\log_{10}\Delta & R \leq R_{o} \\ \mathcal{A}_{o}(T)\log_{10}\Delta & -(R-R_{o})/200, & R > R_{o} \end{cases}$$

with

$$\Delta = S \left(\ln \frac{R^2 + H^2 + S^2}{R^2 + H^2 + S_0^2} \right)^{-\frac{1}{2}}$$

and

$$\Delta_{0} = S \left(\ln \frac{R_{0}^{2} + H^{2} + S^{2}}{R_{0}^{2} + H^{2} + S_{0}^{2}} \right)^{-\frac{1}{2}}, \qquad (1.3.5)$$

where $\mathcal{A}_0(T)$ is an empirically determined parabolic function of T. It is used to calculate the attenuation function at distances R less than R_0 . For distances R > R_0 , the attenuation function is a linear function of distance with slope -1/200. The transition distance R_0 is given by (Model III; Trifunac and Lee, 1985):

$$R_{o} = \frac{1}{2} \left(\frac{-200 \mathcal{A}_{o}(T)(1-S_{o}^{2}/S^{2})}{\ln 10} + \sqrt{\left(\frac{200 \mathcal{A}_{o}(T)(1-S_{o}^{2}/S^{2})}{\ln 10}\right)^{2} - 4H^{2}} \right), \quad (I.3.6)$$

which is a function of H,S (hence M,R,T), S_0 and $\mathcal{A}_0(T)$. Detailed description and plots of $S_{6.5}(T)$, $\mathcal{A}_0(T)$ and R_0 and of the attenuation function $\mathcal{A}tt(\Delta,M,T)$ are all given in Trifunac and Lee (1985).

I.4 THE NEW SCALING RELATION

With the new attenuation function defined, the regression equation of Fourier amplitudes to be used now takes the form:

$$log_{10}FS(T) = M + \mathcal{A}tt(\Delta, M, T) + b_1(T)M$$

$$+ b_2(T)h + b_3(T)v + b_4(T)\Delta/100 + b_5(T) + b_6(T)M^2. \quad (I.4.1)$$

Equation (I.4.1) is of the same form as equation (I.1.2), the old scaling equation, with the old attenuation function $\log_{10} A_0(T)$ replaced by new attenuation function $\mathscr{A}tt(\Delta,M,T)$. The regression analysis is performed on the new database of 1314 components of Fourier amplitude data FS(T), at 91 discrete periods T ranging from 0.04 to 15.0 sec. This is in fact Step 2 of the iteration procedure described in Trifunac and Lee (1985) for the determination of the new attenuation function $\mathscr{A}tt(\Delta,M,T)$, and is identical to the regression analysis procedure used with the old database (Trifunac, 1976a,b; Trifunac and Lee, 1978). For completeness, the details of this step are repeated here.

The data are screened to minimize possible bias in the model that could result from possible uneven distribution of data among the different magnitudes and from excessive contribution to the database from several abundantly recorded earthquakes. To carry out this screening the data are partitioned into six groups corresponding to magnitude ranges: 2.0-2.9, 3.0-3.9, 4.0-4.9, 5.0-5.9, 6.0-6.9 and 7.0-7.9. The data in each of these magnitude ranges are next subdivided according to the site classifications s = 0,1 and 2. The data within each of these subgroups were then divided into 2 sets corresponding to horizontal (v=0) and vertical (v=1) components. The resulting data in each of the groups correspond to the Fourier spectral amplitudes from a

specified earthquake magnitude range for a specified site classification and with specified component orientation. To properly balance the effects of attenuation at small and large distances, the data in each of the subgroups are subdivided further into 2 sets: one for epicentral distances \leq 100 km and the other for distances > 100 km. The data in each of these two final subsets are then arranged in increasing order in terms of their amplitudes. If the number of data points in the first set (R \leq 100 km) is less than 19, all the data points are taken. If there are more than 19 points in this first set, at most 19 points are selected from among the ordered set of data so that they correspond uniformly, as close as possible, to the 5%, 10%, ..., 90% and 95% percentiles at distances $R \le 100$ km. Similarly, at most 5 points are selected from the second set (R > 100 km) of data so that they correspond uniformly to around 16 2/3%, 33 1/3%, 50%, 66 2/3 and 83 1/3% percentiles at distances R > 100 km. This approximate scheme has the effect of reducing the biases described above. Note that this selection process is repeated for each of the 91 periods in the range .04 sec. to 15 sec. At the long period end, the Fourier data whose amplitudes are below that of the average digitization noise i.e., those with signal-to-noise ratio less than one, are automatically eliminated before the above selection process. This will be the case for many of the data from earthquakes of smaller magnitudes and/or recorded at sites of larger epicentral distances. The number of data points used in the regression analysis at the long period end are thus comparatively smaller than those at the rest of the period ranges. The resulting fitted coefficients at each period T resulting from linear

regression will be denoted by $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_4(T)$, $\hat{b}_5(T)$ and $\hat{b}_6(T)$, (equation (I.4.1)) respectively.

I.5 THE NEW REGRESSION COEFFICIENTS

During the regression analysis, it was found that the linear term in \triangle in (I.4.1), $b_4(T)$, is insignificant for most of the periods. Subsequently, this term has been deleted from the regression analysis and the empirical scaling equation, (I.4.1), becomes

$$\log_{10}FS(T) = M + \mathcal{A}tt(\Delta, M, T) + b_{1}(T)M + b_{2}(T)h + b_{3}(T)v + b_{5}(T) + b_{6}(T)M^{2}$$
 (I.5.1)

Figure I.5.1 shows $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_5(T)$ and $\hat{b}_6(T)$ (solid lines) and the estimates of their 80%, 90% and 95% confidence intervals (Westermo and Trifunac, 1978), represented by the corresponding dashed lines.

Substituting these coefficients in equation (I.5.1) gives $\widehat{FS}(T)$, where:

$$\log_{10} \hat{FS}(T) = M + \mathcal{A}tt(\Delta, M, T) + \hat{b}_{1}(T)M + \hat{b}_{2}(T)M + \hat{b}_{3}(T)V + \hat{b}_{5}(T) + \hat{b}_{6}(T)M^{2}, \quad (1.5.2)$$

 $\hat{FS}(T)$ then represents the least squares estimate of the Fourier amplitude spectrum at period T.

For given values of T, h, v and \triangle , $\log_{10} FS(T)$ represents a parabola when plotted versus M. Following the previous work, it is also assumed in the present analysis that equation (I.5.2) applies only in the range $M_{min} \leq M \leq M_{max}$, where $M_{min} = -b_1(T)/(2b_6(T))$ and $M_{max} = -(1+b_1(T))/(2b_6(T))$. Equation (I.5.2) is then modified to:

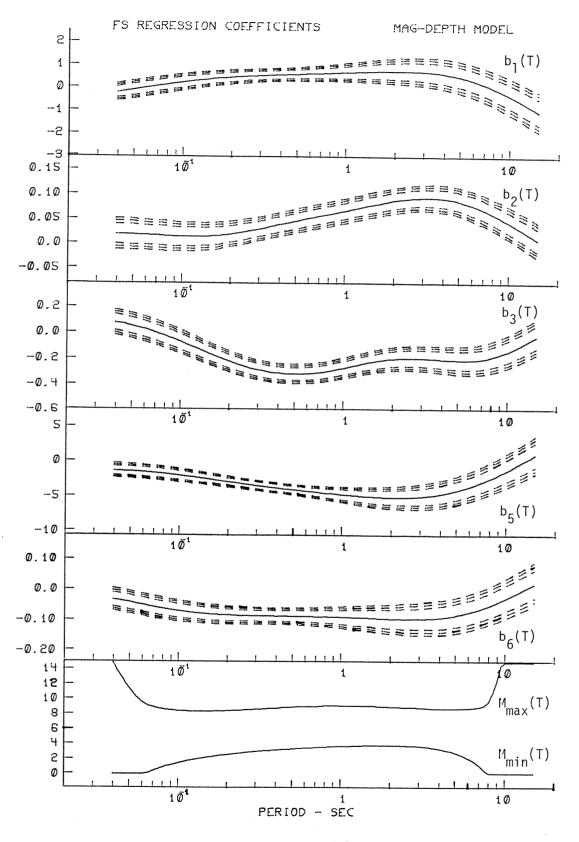


Figure I.5.1

$$\begin{split} \log_{10} \hat{FS}(T) &= \mathscr{A} tt(\Delta, M, T) + \\ \begin{cases} M + \hat{b}_{1}(T) M_{min} + \hat{b}_{2}(T) h + \hat{b}_{3}(T) v + \hat{b}_{5}(T) + \hat{b}_{6}(T) M_{min}^{2}, & M \leq M_{min} \\ M + \hat{b}_{1}(T) M + \hat{b}_{2}(T) h + \hat{b}_{3}(T) v + \hat{b}_{5}(T) + \hat{b}_{6}(T) M^{2}, & M_{min} \leq M \leq M_{max} \\ M_{max} + \hat{b}_{1}(T) M_{max} + \hat{b}_{2}(T) h + \hat{b}_{3}(T) v + \hat{b}_{5}(T) + \hat{b}_{6}(T) M_{max}^{2}, & M_{max} \leq M \end{cases} \end{split}$$

$$(1.5.3)$$

In other words, for M \leq M $_{min}$, M $_{min}$ is used in the terms following M in (I.5.3), i.e. M_{min} is used with $b_1(T)$ and $b_6(T)$. For M \geq M $_{max}$, M_{max} is used in all the terms. This will result in linear growth of $\log_{10} FS(T)$ with M for M \leq M $_{min}$, in parabolic growth for M $_{min} \leq$ M \leq M $_{max}$ and in constant FS(T) corresponding to M $_{max}$ for all M \geq M $_{max}$. The bottom curves of Figure I.5.1 show M $_{min}$ and M $_{max}$ plotted versus T.

With FS(T) representing the Fourier amplitude spectra computed from recorded accelerograms, the residues were calculated as in the previous analysis (Trifunac and Lee, 1978), where the residue given by $\varepsilon(T) = \log_{10}[FS(T)] - \log_{10}[FS(T)], \text{ describes the distribution of the observed FS(T) about the estimated <math>FS(T)$. As in the previous work, it is assumed that $\varepsilon(T)$ can be described by a normal distribution function with mean $\mu(T)$ and standard deviation $\sigma(T)$ as follows:

$$p(\varepsilon,T) = \frac{1}{\sigma(T)\sqrt{2\pi}} \int_{-\infty}^{\varepsilon(T)} \exp\left[-\frac{1}{2} \left(\frac{x-\mu(T)}{\sigma(T)}\right)^{2}\right] dx , \qquad (I.5.4)$$

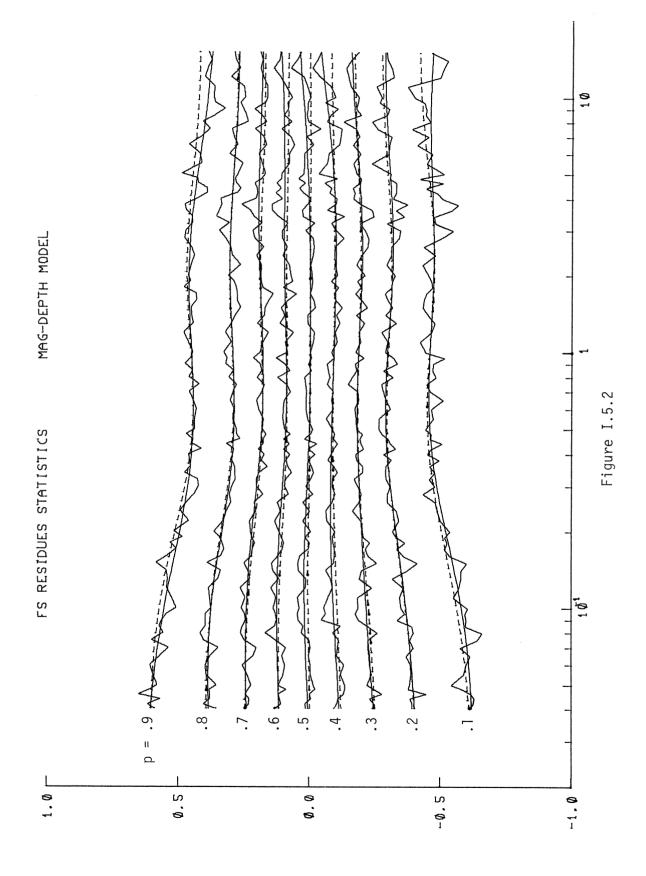
where p(ϵ ,T) represents the probability that $\log_{10}[FS(T)] - \log_{10}[\hat{FS}(T)] \le \epsilon(T)$.

For a given residual value $\varepsilon(T)$ at a particular period T, the actual probability $P^*(\varepsilon,T)$ that $\varepsilon(T)$ will not be exceeded can be evaluated by

finding the fraction of residues $\epsilon(T)$ (computed from the database at that particual period) which are smaller than a given value.

For $p^*(\epsilon,T)$ calculated at 91 periods, $\epsilon(T)$ corresponding to $p^* =$ $0.1, 0.2, \ldots, 0.8$ and 0.9 are plotted in Figure I.5.2. The nine sets of curves, plotted versus period, T, from bottom to top correspond to one at each of the probability levels, 0.1 through 0.9. At each of the nine probability levels, the rough solid curve represents the actual calculated residuals at that particular level. The smooth solid curves are obtained by smoothing the rough solid curves along the T-axis. The smooth surface $p*(\varepsilon,T)$, from the nine smooth solid curves thus represents the distribution of data (FS(T) computed from recorded accelerograms) about the estimate $\widehat{FS}(T)$ in (I.5.1). By fitting $p(\varepsilon,T)$ in (I.5.4) to $p*(\varepsilon,T)$ at 91 periods, the mean and standard deviation of the assumed normal distribution function, respectively $\hat{\mu}(T)$ and $\hat{\sigma}(T)$, can thus be evaluated. Substituting these values into (I.5.4) with $p(\epsilon,T)$ equal to 0.1 through 0.9 will result in $\widehat{\epsilon}(T)$ for the nine probability levels to be calculated. These are the nine dashed lines in Figure I.5.3. The surface $p*(\epsilon,T)$ that resulted from the new model in the present analysis (Figure I.5.2) is narrower in ϵ range when compared to the corresponding surface in our previous analysis (Figure 2 of Trifunac and Lee, 1978).

To test the quality of fit of $\hat{p}(\varepsilon,T)$ to $p*(\varepsilon,T)$ as in the previous analysis, two statistical tests for goodness of fit, namely the Kolmogorov-Smirnov (K-S) and the X^2 tests of the hypothesis that $p*(\varepsilon,T)$ can be approximated by a normal distribution $p(\varepsilon,T)$ (equation (I.5.4)) have been performed. The Kolmogorov-Smirnov statistic calculates, for each period T, the maximum allowed difference between the estimated and the calculated probability levels:



$$KS(T) = \max_{\varepsilon} |p(\hat{\varepsilon},T) - p*(\varepsilon,T)|. \qquad (I.5.5)$$

To calculate the X^2 statistic, a standard procedure was adopted, which is different from that of the previous analysis, so that the number of intervals used is the same for all periods, T:

Step 1: The ε -axis of residues is subdivided into 10 intervals I_0 , I_1,\ldots,I_9 such that $I_k=\left[\underline{\varepsilon}_k,\underline{\varepsilon}_{k+1}\right]$ for each i contains residues in the probability levels between p_k and p_{k+1} , where $p_k=0.1k$. The residues are divided in intervals of 10% probability each: 0-10%, 10-20%, ... to 90-100%. Note that $\underline{\varepsilon}_0=-\infty$ and $\underline{\varepsilon}_{10}=+\infty$.

Step 2: For each $\underline{\varepsilon}_k$, the estimated probability level $\hat{p}_k = \hat{p}(\underline{\varepsilon}_k, T)$ is computed from (I.5.4) using the estimated coefficients $\hat{\mu}_k(T)$ and $\hat{\sigma}_k(T)$. Note that $\hat{p}_0 = \hat{p}(\underline{\varepsilon}_0, T) = 0$ and $\hat{p}_{10} = \hat{p}(\underline{\varepsilon}_{10}, T) = 1$. The estimated probability that the residue ε assumes any value in the interval $I_k[\underline{\varepsilon}_k, \underline{\varepsilon}_{k+1}]$, for k = 0 to 9, is then given by the difference $\hat{p}_{k+1} - \hat{p}_k$. The actual probability that the residue ε assumes value in any of the intervals is of course the chosen value of .1 (10%).

Step 3: In each interval I_k , k = 0 to 9, define

$$\hat{n}_{k} = (\hat{p}_{k+1} - \hat{p}_{k})N$$
, and $n_{k} = 0.1 N$, (I.5.6)

where N is the total number of residues in all the intervals. $\hat{n_k}$ is the estimated number of residuals theoretically expected in the interval I_k and n_k is the actual number of residues in the same interval. The χ^2 statistics are then calculated from the formula

$$X^{2}(T) = \sum_{k=1}^{9} \frac{(\hat{n}_{k} - n_{k})^{2}}{\hat{n}_{k}} \qquad (1.5.7)$$

Another convenient form of (I.5.7) for χ^2 is

$$X^{2}(T) = N \sum_{k=0}^{9} \frac{(\hat{p}_{k+1} - \hat{p}_{k} - .1)^{2}}{(\hat{p}_{k+1} - \hat{p}_{k})} . \qquad (I.5.8)$$

Note that this differs from the $\rm X^2$ statistics used in the previous analysis by having the factor N. The above definition is adopted from Kreyszig (1972).

Having computed the values of KS(T) and χ^2 (T) for each period T, those were compared with their corresponding cutoff values at 95% probability level. For the Kolmogorov-Smirnov statistic,

$$P(KS \le C) = 95\% \Rightarrow C = 0.058.$$
 (I.5.9)

Thus if KS(T) < 0.058, the hypothesis that the probability distribution is normal will not be rejected. Similarly, for the x^2 -statistic, with K = 10 intervals, r = 2 the number of parameters used in estimating \hat{n}_k , the number of degrees of freedom is K-r-l = 7, and

$$P(X^2 \le C) = 95\% \implies C = 14.07.$$
 (I.5.10)

Thus again if $\chi^2(T)$ < 14.07, the hypothesis that the probability distribution is normal will not be rejected.

Figure I.5.3 shows a plot of the statistical parameters of the residuals. The smooth amplitudes of $\hat{\mu}(T)$ and $\hat{\sigma}(T)$ and their 95% confidence intervals are given in the top 2 plots of the figure respectively. The two full curves in the bottom of Figure I.5.3 show the smoothed amplitudes of the computed $X^2(T)$ and KS(T) respectively. The dashed

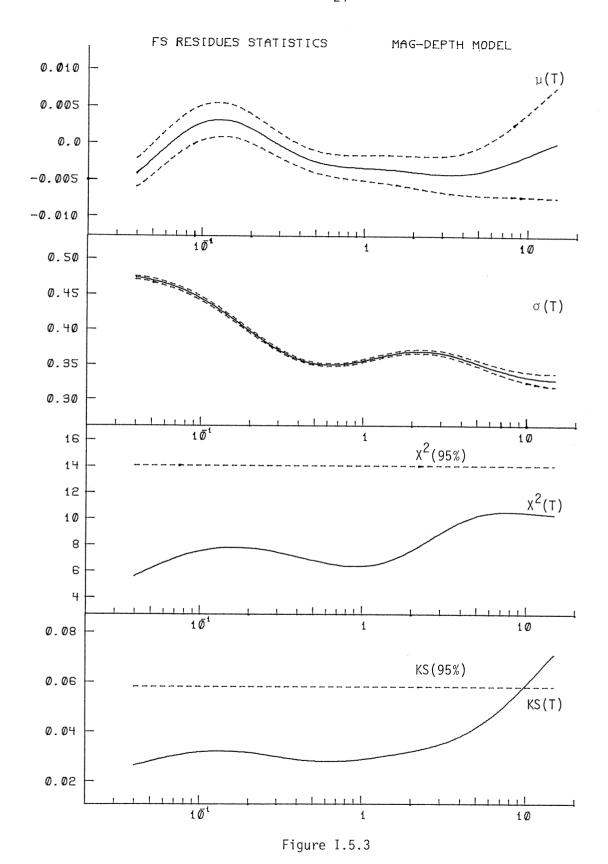


TABLE I.5.1

$\log_{10}FS(T) = M + \mathcal{A}tt(\Delta,M,T) + b_1(T)M + b_2(T)h + b_3(T)v + b_5(T) + b_6(T)$)M2
$= M + \mathcal{A}tt(\Delta, M, T) + b_1(T)M + b_2(T)h + b_3(T)v + b_5(T)$	
= $M + \mathcal{A}tt(\Delta, M, T) + b_1(T)M + b_2(T)h + b_3(T)v +$	E
$= M + \mathcal{A}tt(\Delta, M, T) + b_1(T)M + b_2(T)h + b_3(T)$	+ >
$= M + \mathcal{A}tt(\triangle,M,T) + b_1(T)M + b_2($	b ₃ (
$= M + \mathcal{A}tt(\triangle,M,T) + b_1(T)M +$	
$= M + \mathcal{A}tt(\triangle,M,T) +$	+
$= M + \mathcal{A}tt(\Delta, M)$	+ b ₁ (T
11	øtt(∆,M
log ₁₀ FS(T)	11
	log ₁₀ FS(T)

14.00	933 013 050 556 .011 .000	468 287 160 045 .037 .109 .182 .271	000 .328 10.308
7.50	.048 .063 196 -2.698 058 .416	461 296 177 076 .016 .099 .189	003 .338 10.563
4.40	.549 .089 -214 -4.477 -090 3.038 8.573	467 310 193 095 097 196 294	004 .356 9.990
2.80	. 718 . 093 - 196 -5.141 - 098 3.656 8.748	474 318 201 001 .097 .195	004 .367 8.689
1.60	.734 .082 219 -5.187 096 3.811	470 313 197 099 003 .187 .303	004 .365 7.007
06•	. 668 . 063 - 287 -4.766 - 092 3.640 9.090	458 296 187 093 004 . 088 . 178 . 291	003 .353 6.365 .029
• 50	.594 .045 325 -4.164 090 3.305	461 293 184 087 000 .090 .182 .288	003 .351 6.813 .028
• 34	. 545 . 033 . 308 -3.724 - 089 3.053 8.654	479 305 084 .006 .098 .195	001 .364 7.295 .029
.19	.432 .017 219 -2.967 086 2.508 8.308	527 336 203 084 .015 .114 .220	.002 .401 7.795
.11	. 251 . 012 - 098 -2.241 - 076 1.660 8.274	575 366 217 092 . 122 . 237 . 368	.003 .438 7.671 .032
• 065	. 014 . 014 . 007 -1.714 056 . 126 8. 992	607 389 232 105 .012 .120 .242 .384	.000 .462 6.823
.040	228 . 018 . 077 -1. 401 034 . 000	621 404 246 115 .004 .117 .241	TCS:004473 5.655
PERIOD, T (SEC)	COEFFICIENTS: b1(T) b2(T) b3(T) b3(T) b5(T) Mmin	ES.	E STATISTICS:
PERIOL	COEFF1 b1(T) b2(T) b3(T) b6(T) Mmin max	RESIDUES: p = .1 p = .2 p = .3 p = .4 p = .5 p = .6 p = .7 p = .8	RESIDUE µ(T) σ(T) X ² (T) KS(T)

lines are their corresponding 95% cutoff levels. It is seen that with the minor exception in the long period end, both the X^2 and K-S tests fail to reject the hypothesis that the distribution is normal. The function $p(\varepsilon,T)$ in (I.5.4) thus represents an acceptable approximation to $p*(\varepsilon,T)$.

Table I.5.1 presents, for 12 periods, between T = 0.04 sec and T = 14 sec, the amplitudes of the smoothed regression coefficients $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_5(T)$, $\hat{b}_6(T)$ (note that $b_4(T)$ has been deleted), $\hat{M}_{min}(T)$, $\hat{M}_{max}(T)$, the nine smoothed calculated residue levels corresponding to p*(ϵ ,T) = 0.1, 0.2,...,0.8 and 0.9, the smoothed amplitudes $\hat{\mu}(T)$, $\hat{\sigma}(T)$ in equation (I.5.4), the X² and the Kolmogorov-Smirnov statistics. The 12 periods used appear to be sufficient for most practical computations since the smoothness of the coefficients is such that almost any interpolation scheme will yield adequate estimates of the FS(T) amplitudes at any period in the range between 0.04 sec and 15 sec.

I.6 EXAMPLES OF ESTIMATED FOURIER SPECTRA

Figures I.6.1 and I.6.2 present examples of FS(T) computed for M = 4.5, 5.5, 6.5, 7.5 at R = 0, H = 5 km, or then $\Delta \approx 5$ km, for $p(\epsilon,T)=0.5$, and the spectral amplitudes which have signal-to-noise ratio greater than one (Trifunac, 1976a). Figure I.6.1 is for horizontal motion while Figure I.6.2 is for vertical motion. The solid lines in both figures correspond to an alluvial depth of h = 0 km while the dashed lines correspond to h = 4 km.

The trends of computed FS(T) spectral amplitudes in the figures are in many ways similar to those discussed by Trifunac and Lee (1978). The rate of growth of amplitudes with earthquake magnitude, M clearly decreases as M approaches M = 7.5. The effect of the depth of sediments beneath the recording station is important only for intermediate to long periods and is small at short periods.

Comparison of the two figures shows that the vertical spectral amplitudes are smaller than the horizontal spectral amplitudes, except at short periods.

Figure I.6.3 illustrates the effects of epicentral distance R on the changes of spectral amplitudes for $p(\varepsilon,T)=0.5$, magnitude M = 6.5, focal depth H = 5 km, sedimentary depth h = 2 km and for horizontal (solid lines) and vertical (dashed lines) components. Four sets of curves corresponding to R = 0, 25, 50 and 100 km are shown. As can be seen from equation (I.5.1), the only term governing these distance changes is the change in Δ , the representative distance, in the frequency dependent attenuation term $\mathscr{A}tt(\Delta,M,T)$. As already noted in Trifunac and Lee (1985), the attenuation of the high frequency waves is somewhat faster than that of the low-frequency waves.

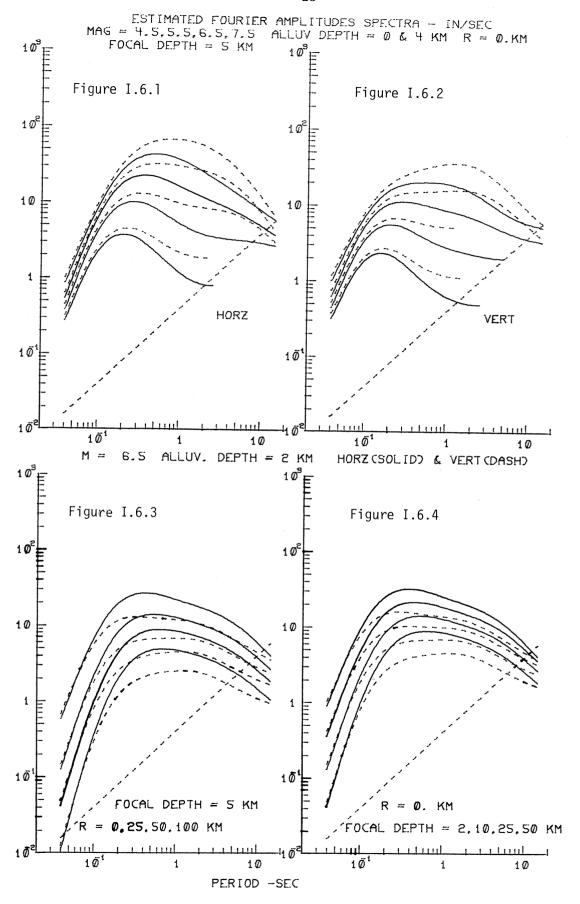


Figure I.6.4 illustrates the effects of focal depth H on the changes of spectral amplitudes for $p(\epsilon,T)=0.5$, M=6.5, R=0 km, h=2 km and for horizontal (solid lines) and vertical (dashed line) components. Four sets of curves corresponding to H=2, 5, 25 and 50 km are presented. As in the case for distances R, the attenuation of the high frequency waves with focal depth is faster than that of the low frequency waves.

Figures I.6.5 and I.6.6 show an example of how horizontal and vertical Fourier spectra computed from equations (I.5.2) and (I.5.4) compare with the acceleration spectra for the three components of strong-motion recorded in El Centro during the Imperial Valley, California earthquake of 1940. During this earthquake, the fault rupture was initiated most probably at a distance of about 10 km, southeast of El Centro. With the introduction of fault size and focal depth now available in the new scaling model, a reduction in the observed differences between the computed and estimated Fourier spectra is to be expected. In these figures, as in the previous analysis, the $log_{10}FS(T)$ spectra were computed for p(ϵ ,T) = 0.1, 0.5 and 0.9. The interval between the spectra for p = 0.1 and 0.9 represents an estimate of the 80% confidence interval. The statistical parameters used are M = 6.4, R = 9.3 km, focal depth H = 5 km and alluvial depth h = 15000 ft. As may be concluded from these figures, the agreement between the recorded and empirically predicted spectra in this case is very satisfactory, as in the previous analysis. The only difference between the present and the previous analysis is that the 80% confidence interval is now narrower than before.

AA001 EL CENTRO, 1940 COMP HORZ $M = 6.4 \quad R = 9.3 \text{KM} \quad \text{FH} = 5.0 \text{KM} \quad \text{DEPTH} = 15000.FT \quad \text{V} = 0.$

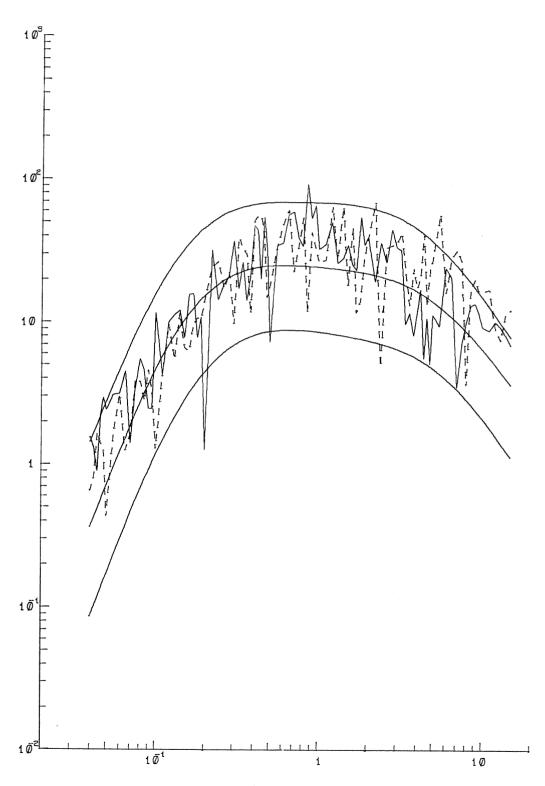


Figure I.6.5

AA001 EL CENTRO, 1940 COMP VERT M = 6.4 R = 9.3KM FH = 5.0KM DEPTH =15000.FT V = 1.

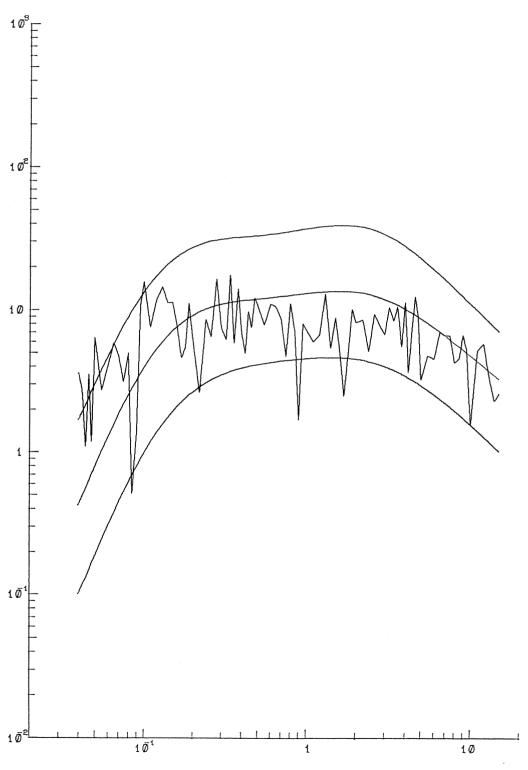


Figure I.6.6

AC041 PAC0IMA DAM, 1971 COMP HORZ M = 6.4 R = .0KM FH = 2.0KM DEPTH = 0.FT V = 0.

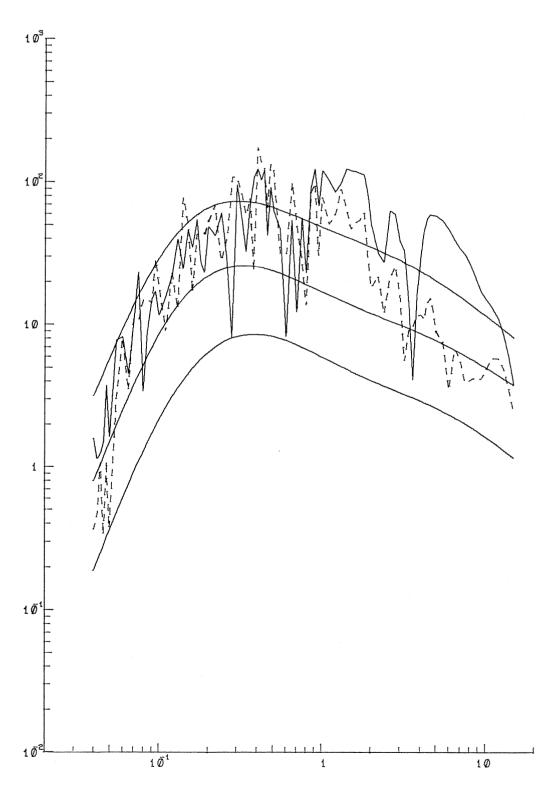


Figure I.6.7

AC041 PAC0IMA DAM, 1970 COMP DOWN $M=6.4\ R=.0KM\ FH=2.0KM\ DEPTH=0.FT\ V=1.$

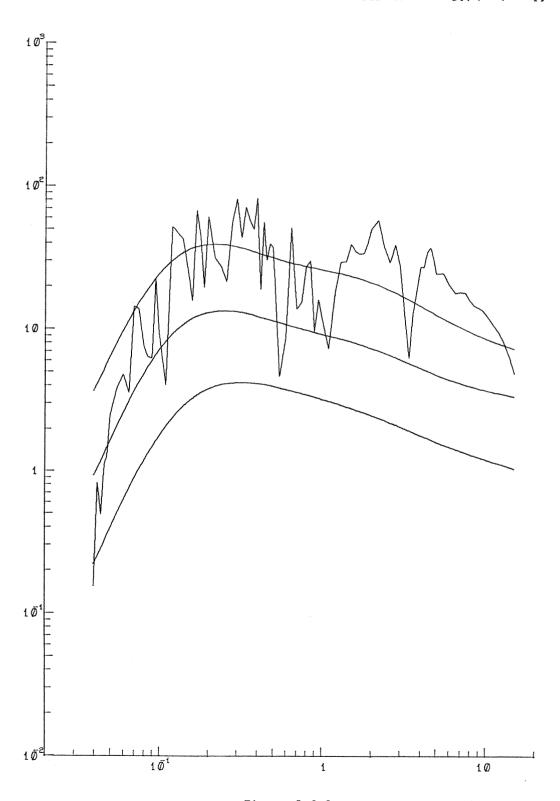


Figure I.6.8

This type of agreement between empirically predicted and actually recorded spectra may be expected in many of the average or better than average cases. An example of a fit which is below average is illustrated in Figures I.6.7 and I.6.8 for the spectra of strong-motion accelerograms recorded at the Pacoima Dam site during the San Fernando, California earthquake of February 9, 1971. During this earthquake the fault rupture passed underneath the dam site (so R = 0 km) at a "focal depth" of H \cong 2 km. The other scaling parameters used are M = 6.4 and h = 0 ft. The average trend of the spectra of the recorded accelerograms is larger than that of the 50th percentile (p = 0.5) of the estimated spectra. Nevertheless, the estimated curves do follow the overall amplitude and shape trends quite well.

PART II: SCALING OF FOURIER SPECTRA IN TERMS OF M, R, S, s AND V

II.1 THE SCALING RELATION

Part II of this work continues the description of the preliminary empirical model for scaling Fourier amplitude spectra of strong ground motion in terms of earthquake magnitude, source-to-station "representative" distance and the local site geology. Part I characterizes the local geology by the approximate overall depth of sedimentary deposits beneath the recording station, h, in km. It has been pointed out previously, that while the depth of sediments at each recording station represents a preferable site characterization, in many instances, little may be known about the depth of alluvium and sedimentary layers at some sites so that the scaling of amplitudes at such sites using depth, h, is not possible. The site characterization in terms of s = 0, l and 2 (Trifunac and Brady, 1975), which can be determined from knowledge of surface geology only, thus remains a useful approach to the scaling of strong-motion amplitudes.

Ideally, a site should be classified either as an alluvium site (s=0) or as a basement rock site (s=2). The intermediate classification (s=1) should be used only in complicated cases when it is not obvious how to select s=0 or s=2, typically corresponding to consolidated sedimentary rock or to a complex geologic environment (Trifunac and Brady, 1975), or when information on the site geology is not available. With this classification, of the 186 records of the original database, 87 records (63%) have been recorded on alluvium sites, 43 records (23%) come from "intermediate" sites and only 15 records (8%) were recorded on basement rock sites. With the new

database of 438 records now available 313 records (71%) have been recorded on alluvium sites (s=0), 76 records (17%) on "intermediate" sites (s=1) and 47 records (11%) on basement rock sites (s=2). The majority of records in the new database thus continue to come from alluvium sites.

Following Part I of this report, and adopting the form of the scaling relation in equation (I.5.1), the required scaling relation now takes the form

$$log_{10}FS(T) = M + \mathcal{A}tt(\Delta,M,T) + b_1(T)M + b_2(T)s + b_3(T)v + b_5(T) + b_6(T)M^2, \quad (II.1.1)$$

with all parameters defined as before. $b_2(T)$ is now the coefficient associated with the site condition s. Note that the term $b_4(T) \triangle / 100$ is again deleted.

The scaling functions $b_1(t)$ through $b_6(T)$ are determined through the regression analysis of the new database of 1314 components of spectral amplitudes, FS(T) at 91 discrete periods T ranging from 0.04 sec to 15.0 sec. As in Part I of this report, the data are first screened for possible bias in the model. All procedures in data preparation and selection, and the form of the regression analysis employed here are identical to those in Part I of this report, and their description need not be repeated here.

The resulting fitted coefficients, at each period T, from linear regression are again denoted by $\hat{b}_1(T)$ through $\hat{b}_6(T)$, (as in equation (I.5.2)), respectively. Much of the format of the description in this and the sections to follow will almost be identical to that in Part I of this work.

Substituting the fitted coefficients in equation (II.1.1) gives $\hat{\mathsf{FS}}(\mathsf{T}),$ where

$$\log_{10} \hat{FS}(T) = M + \mathcal{A}tt(\Delta, M, T) + \hat{b}_{1}(T)M + \hat{b}_{2}(T)s + \hat{b}_{3}(T)v + \hat{b}_{5}(T) + \hat{b}_{6}(T)M^{2}$$
 (II.1.2)

As before, equation (II.1.2) applies only in the range $\rm M_{min} \leq \rm M \leq M_{max},$ where

$$M_{\min} = -\hat{b}_1(T)/(2\hat{b}_6(T))$$
 and
 $M_{\max} = -(1 + \hat{b}_1(T))/(2\hat{b}_6(T)),$ (II.1.3)

and equation (II.1.2) is then modified to:

$$\begin{split} &\log_{10}\hat{FS}(T) = \mathscr{A}tt(\Delta,M,T) \ + \\ &\left\{ M + b_1(T)M_{\text{min}} + \hat{b}_2(T)s + \hat{b}_3(T)v + \hat{b}_5(T) + \hat{b}_6(T)M_{\text{min}}^2, \quad M \leq M_{\text{min}} \\ M + \hat{b}_1(T)M + \hat{b}_2(T)s + \hat{b}_3(T)v + \hat{b}_5(T) + \hat{b}_6(T)M^2, \quad M_{\text{min}} \leq M \leq M_{\text{max}} \\ M_{\text{max}} + \hat{b}_1(T)M_{\text{max}} + \hat{b}_2(T)s + \hat{b}_3(T)v + \hat{b}_5(T) + \hat{b}_2(T)M_{\text{max}}^2, \quad M_{\text{max}} \leq M \end{split} \right.$$

The residues $\varepsilon(T) = \log_{10}[FS(T)] - \log_{10}[\widehat{FS}(T)]$ describing the distribution of the observed FS(T) about the estimated $\widehat{FS}(T)$ are next calculated. As in the previous part, $\varepsilon(T)$ is described by a normal distribution function with mean $\mu(T)$ and standard deviation $\sigma(T)$.

II.2 THE REGRESSION COEFFICIENTS

Figure II.2.1 shows the smoothed coefficients $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_5(T)$ and $\hat{b}_6(T)$ (solid lines) together with the estimates of their 80%, 90% and 95% confidence intervals (dashed lines). Comparison of this figure with the corresponding Figure I.5.1 in the previous Part I of this work shows that the functions $\hat{b}_1(T)$, $\hat{b}_3(T)$, $\hat{b}_5(T)$ and $\hat{b}_6(T)$ as given respectively by the first, third, fourth and fifth graphs from the top are almost identical. The functions correspond to the same parameters, M, v, 1 and M^2 , respectively, in the scaling relations and their similarity demonstrates the consistency of the two models of scaling. The functions $\hat{b}_2(T)$ as given by the second graph from the top in both figures are of opposite sign, which again is consistent with the models since s = 2 corresponds to h = 0 km (basement rock), while s = 0 corresponds to h >> 0 km (alluvium). The bottom graph of Figure II.2.1 shows M_{min} and M_{max} as given by equation (II.1.3).

Figure II.2.2 shows the plot of the residual levels corresponding to p* (ε ,T) = 0.1, 0.2,...,0.8 and 0.9 for \log_{10} FS(T). Refer to the same Figure I.5.2 in Part I of this work for a complete description of each of the nine sets of curves. It is of interest to compare these two figures, since those illustrate the spread of the observed data about their corresponding models, which differ only in the characterization of local site geology. The resemblance of the two figures demonstrates clearly that the uncertainties associated with the characterization of local geology in terms of site parameters s = 0, 1 and 2 are not much greater than those associated with site characterization in terms of alluvial depth, h.

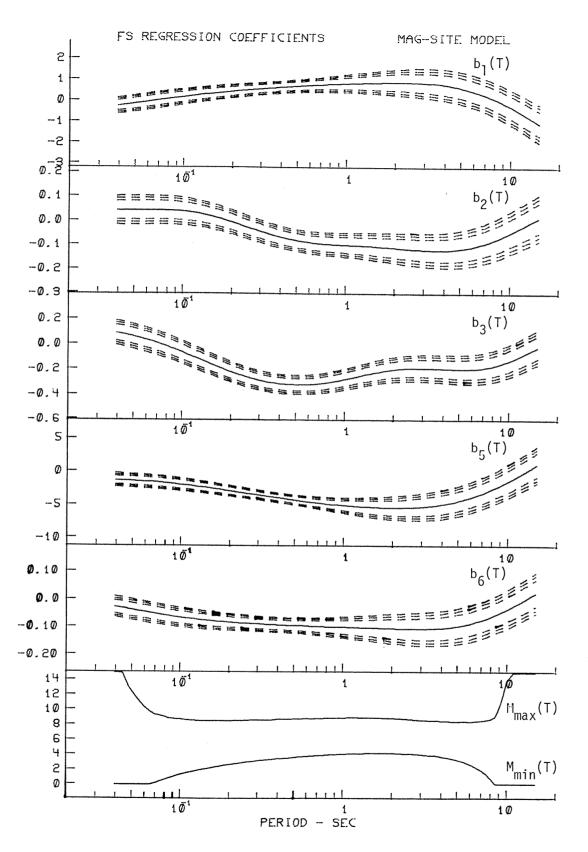
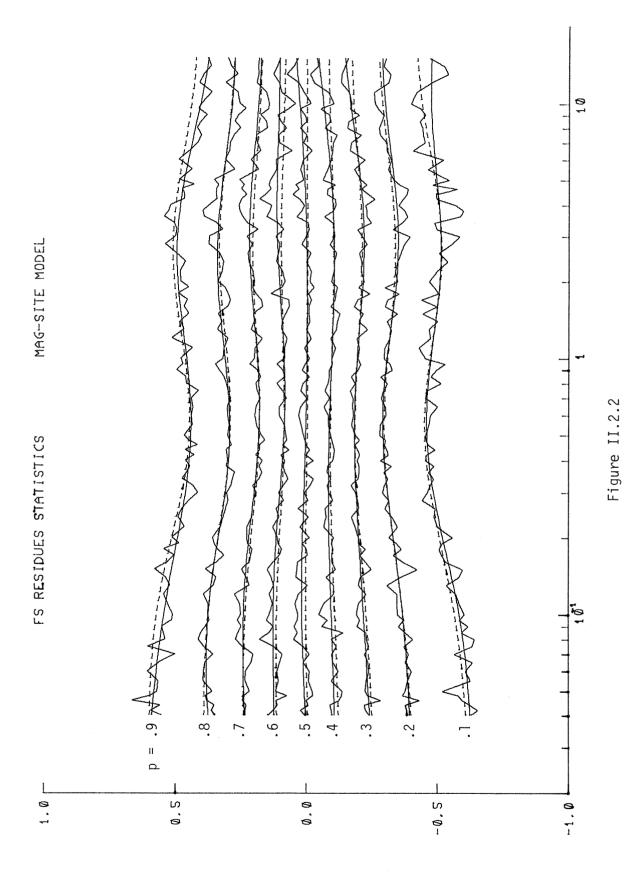


Figure II.2.1



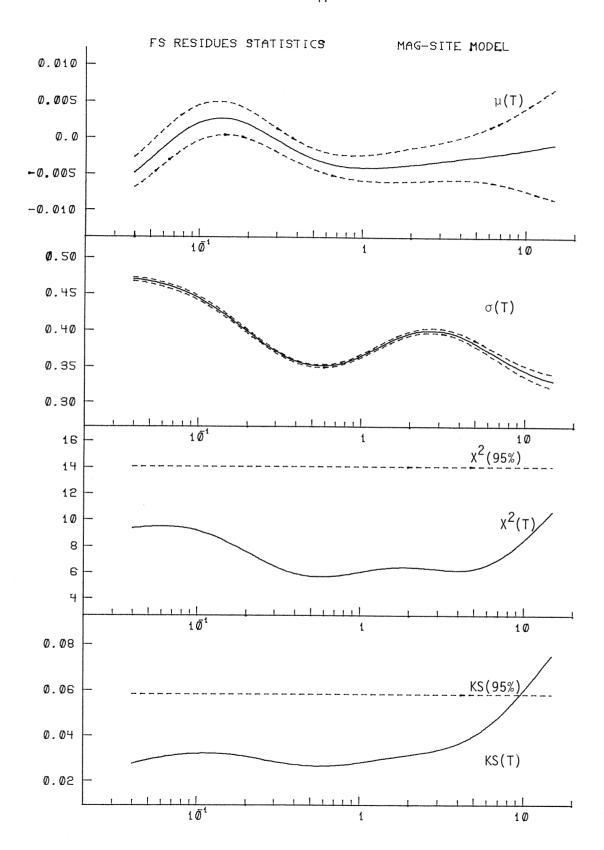


Figure II.2.3

TABLE II.2.1

	14.00	933 . 002 044 . 580 . 012 . 000	475 290 150 048 .038 .104 .175	001 .331 10.267 .073
	7.50	. 184 086 183 -2. 924 069 1. 340 8. 608	481 313 184 078 . 019 . 109 . 193	002 .358 7.257 .050
T)M ²	4.40	. 712 121 201 -4. 741 102 3. 478 8. 361	502 341 211 098 .005 .114 .213	003 .388 6.194 .038
$T) + b_6(T)M^2$	2.80	.869 122 191 -5.395 109 3.976 8.551	513 350 219 105 001 . 111 . 218 . 340	003 .400 6.270 .033
$(x + b_5(T))$	1,60	.883 110 223 -5.487 107 4.137 8.823	500 333 209 104 003 . 097 . 204 . 332	004 .388 6.418
$s + b_3(T)v$	06•	. 820 102 294 -5. 100 103 3. 985 8. 845	469 303 191 095 001 .084 .184	004 .361 5.950 .028
$+ b_2(T)s$	• 50	. 706 - 084 - 329 -4.394 - 098 3.598 8.692	460 293 184 088 .003 .086	003 .351 5.687 .027
+ b ₁ (T)M	.34	. 610 057 309 -3. 844 094 3. 240 8. 550	476 305 188 086 .007 .194	001 .363 6.169
$\mathscr{A}tt(\vartriangle,M,T) + b_1(T)M$	•19	.433 003 216 -2.965 086 2.530 8.374	526 337 202 088 014 . 115 . 220 . 337	.002 .402 7.698
M + Att	• 11	. 222 . 033 092 -2. 207 072 1. 542 8. 476	576 365 217 095 .017 .125 .237 .369	.002 .438 9.017
log ₁₀ FS(T) = N	• 065	019 .042 .015 -1.681 052 .000	609 384 231 103 . 012 . 125 . 240 . 380	001 .460 9.498
10910	.040	258 . 041 . 086 -1. 373 030 . 000	624 395 243 108 . 004 . 123 . 374 . 589	.cs:005 .471 9.379
	PERIOD, T (SEC)	COEFFICIENTS: b1(T) b2(T) b3(T) b5(T) b6(T) Mmin max	RESIDUES: p = .1 p = .3 p = .4 p = .5 p = .6 p = .7 p = .8	RESIDUE STATISTICS: µ(T) α(T) X2 KS(T) 9.

Figure II.2.3 shows the plot of the statistical parameters in the description of the residues, namely, $\hat{\mu}(T)$, $\hat{\sigma}(T)$, $\chi^2(T)$ and KS(T) from top to bottom. Note that except for T beyond 10 sec, both the χ^2 and KS tests again fail to reject the hypothesis that the distribution is normal.

Table II.2.1 gives, for 12 periods between T = 0.04 sec and and T = 14 sec, the five regression coefficients, $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_5(T)$ and $\hat{b}_6(T)$, $\hat{M}_{min}(T)$, $\hat{M}_{max}(T)$, the nine residue levels corresponding to p*(ϵ ,T) = 0.1 through 0.9, the coefficients $\hat{\mu}(T)$, $\hat{\sigma}(T)$ of the normal distribution and finally the X²(T) and KS(T) statistics.

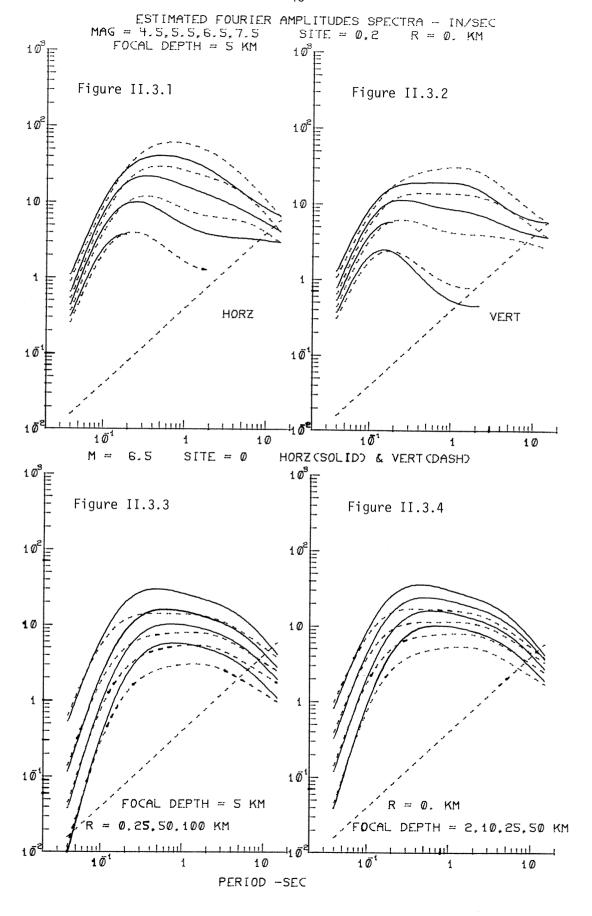
II.3 EXAMPLES OF ESTIMATED FOURIER SPECTRA

Figures II.3.1 and II.3.2 present examples of FS(T) computed for M = 4.5, 5.5, 6.5 and 7.5 at R = 0, H = 5 km for $p(\epsilon,T) = 0.5$. Figure II.3.1 is for horizontal motion (v = 0) while Figure II.3.2 is for vertical motion (v = 1). The solid lines in the figures correspond to s = 2, while the dashed lines correspond to s = 0. The diagonal dashed lines again represent the empirical average Fourier amplitudes of digitization noise.

Comparison of these figures with the corresponding Figures I.6.1 and I.6.2 of Part I shows an interesting trend. It has been noted previously, that for the same magnitude, M, the spectral amplitudes are slightly higher for basement rock (s=2) then for alluvial site (s=0) for periods up to 0.2 seconds, beyond which the trend is reversed. This trend is observed here, but not for the model using depth of sediments as shown in Figures I.6.1 and I.6.2. There it is only observed that the effect of the depth of sediments is negligible at short periods, and for the same magnitude, M, and spectral amplitudes are higher for alluvial site (h >> 0) than for basement roch (h=0) only for intermediate to long periods.

Figure II.3.3 illustrates the effects of epicentral distance R, and Figure II.3.4 shows the effects of focal depth H, on the changes of spectral amplitudes. Comparison of these figures with the corresponding Figures I.6.3 and I.6.4 in Part I of this work shows considerable degree of similarity.

Figures II.3.5 and II.3.6 compare the amplitudes computed from the accelerations recorded in 1940 in El Centro, with those calculated



AA001 EL CENTRO, 1940 COMP HORZ $M = 6.4 \quad R = 9.3 \text{KM} \quad \text{FH} = 5.0 \text{KM} \quad \text{SITE} = 0. \quad \text{V} = 0.$

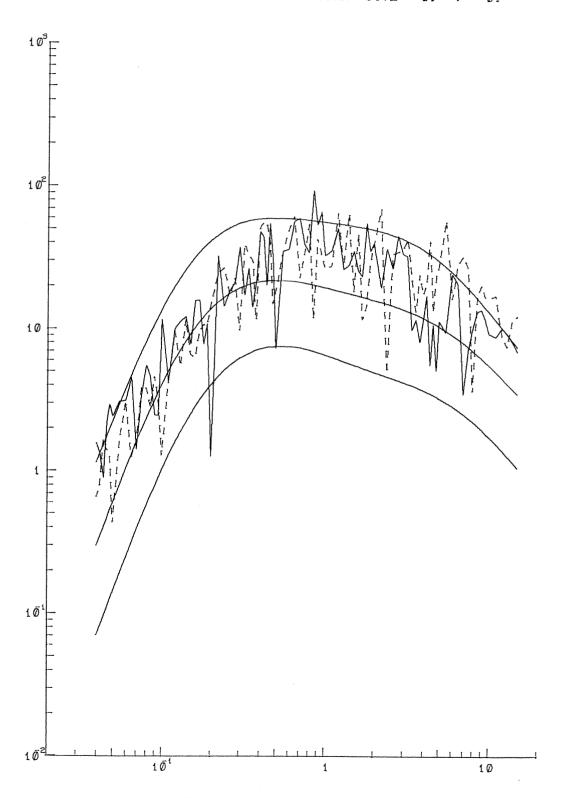


Figure II.3.5

AA001 EL CENTRO, 1940 COMP VERT M = 6.4 R = 9.3KM FH = 5.0KM SITE = 0. V = 1.

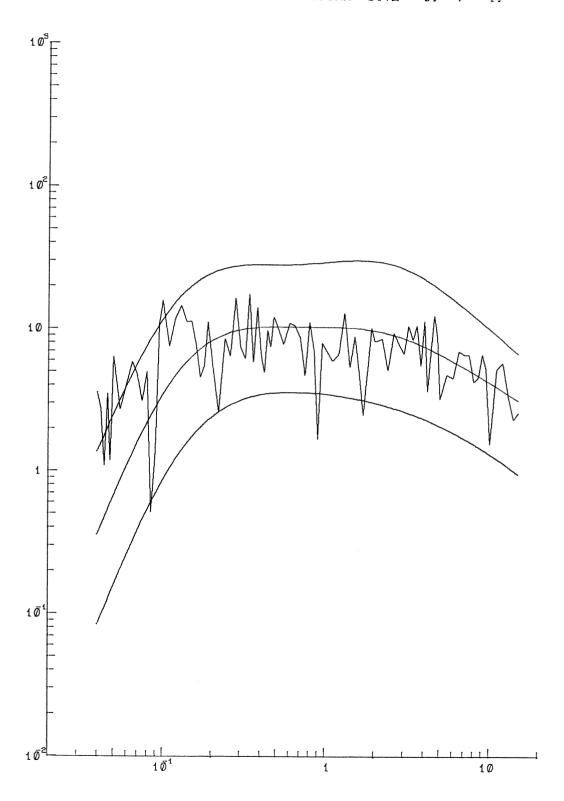


Figure II.3.6

AC041 PACOIMA DAM, 1971 COMP HORZ $M = 6.4 \quad R = .0 \text{KM} \quad \text{FH} = 2.0 \text{KM} \quad \text{SITE} = 2. \quad \text{V} = 0.$

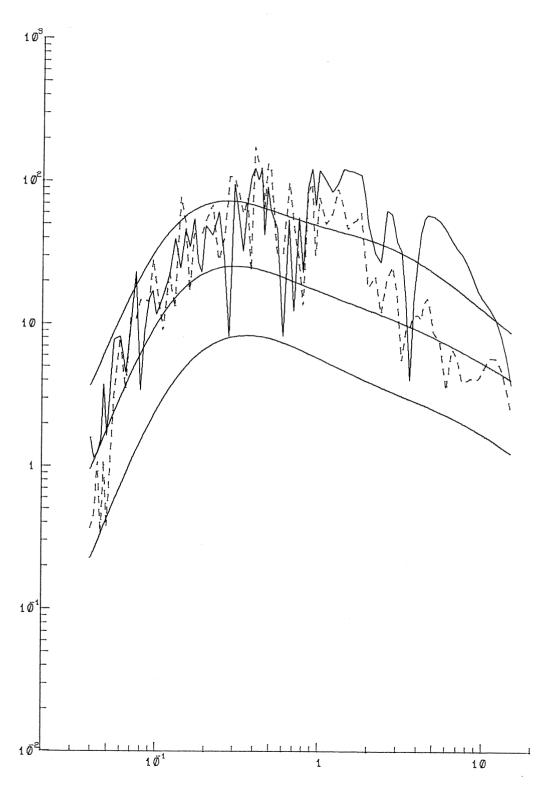


Figure II.3.7

AC041 PAC0IMA DAM, 1971 COMP DOWN $M = 6.4 \quad R = .0 \text{KM} \quad \text{FH} = 2.0 \text{KM} \quad \text{SITE} = 2. \quad \text{V} = 1.$

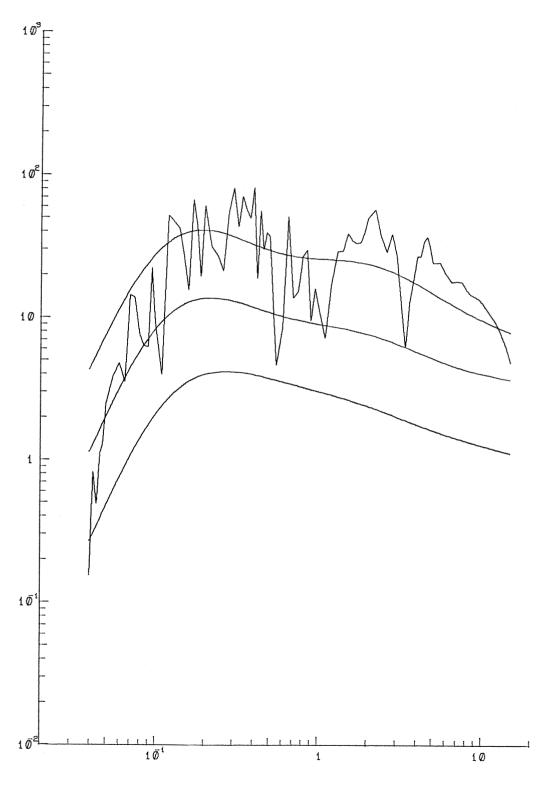


Figure II.3.8

from M = 6.4, R = 9.3 km, H = 5 km and s = 0. The agreement is fair to good. Figures II.3.7 and II.3.8 compare the spectral amplitudes computed from the 1971 Pacoima Dam accelerations with the spectral estimates calculated for M = 6.4, R = 0 km, H = 2 km and s = 2. The fit is again of similar quality, as that in the case of the model in Part I (Figures I.6.7 and I.6.8).

This completes the presentation of Part II of the scaling of FS(T) in terms of M, R, H, s and v.

PART III: SCALING OF FOURIER SPECTRA IN TERMS OF MMI, h AND v

III.1 THE SCALING RELATION

Parts I and II of this work presented a description of the preliminary empirical model for scaling Fourier amplitude spectra of strong earthquake ground motion in terms of earthquake magnitude, souce-to-station distance and a characterization of local geology at the recording station. Parts III and IV of this work will extend this method to the scaling of Fourier amplitude spectra in terms of the Modified Mercalli Intensity (MMI) and a characterization of local geology at the recording site.

Such an analysis was initiated by Trifunac (1979) who pointed out that a description of the expected levels of ground shaking at a site in terms of earthquake intensity will likely remain a common engineering tool for many years to come. Although new instrumentation is now being deployed in many parts of the world, the long-term historical seismicity records are still available only in terms of the locally developed intensity scales. For this reason, it is worthwhile now to re-evaluate the nature of correlations that exist between earthquake intensity and the Fourier amplitudes of recorded ground motions, in spite of all the shortcomings associated with this qualitative and descriptive scaling of strong earthquake ground motion. Further it is worthwhile to re-examine the meaning of such scaling as it still represents an important basis for extrapolating the possible future earthquake risk in many seismic zones of the world.

Trifunac (1979) developed a direct correlation between the Fourier spectrum amplitudes and the reported MMI at the recording stations

using the database of the same 57 earthquakes and the corresponding 186 records that were used earlier (Trifunac (1976b)) for scaling of spectra in terms of earthquake magnitudes. The distribution of these data with respect to the levels of Modified Mercalli Intensity reported at the recording sites is as follows: MMI = III, 1 record; MMI = IV, 3 records; MMI = V, 34 records: MMI = VI, 66 records; MMI = VII, 75 records; MMI = VIII, 6 records, and MMI = X, 1 record. The local geology was then characterized by the geologic site conditions, s, (see Part II of this report for discussion and references on parameters).

Trifunac and Lee (1978) extended this type of analysis by including the depth of sedimentary deposits beneath the recording station, h, as a site characteristic, in place of the site parameter s. The scaling equation then became (equation (5) of Trifunac and Lee (1978)):

$$log_{10}[FS(T)] = b(T)I_{MM} + c(T) + d(T)h + e(T)v$$
, (III.1.1)

with I_{MM} denoting the reported levels on the MMI scale, and with all other parameters defined as above. The form of equation (III.1.1) was chosen on the basis of several previous analyses (Trifunac (1976a,b); Trifunac and Anderson (1977)), suggesting that only the linear terms in the equations are statistically significant and that all higher order terms as well as the mixed terms may be omitted from the analysis.

The same analysis has been carried out now on the new database of 438 free-field records from 104 earthquakes. The new set of earthquakes that have been added to the catalogue through the years 1972 to 1981 have magnitudes typically below 6. For this reason the MMI levels for many of these earthquakes are not well documented or have not been reported. One possible approach then to estimating spectral amplitudes

of strong shaking in terms of MMI at the recording site would be first to develop correlations between the MMI levels with the corresponding earthquake magnitude, representative source to station distance and local site geology. Lee and Trifunac (1985) used the original database of 57 earthquakes and 186 stations where the reported MMI levels are available and performed the above correlation with the following equation:

$$I_{MM} = 1.5M - A - B \ln \Delta - C \Delta / 100 - Ds$$
 (III.1.2)

where the parameters M, \triangle and s are defined as before. The estimated MMI levels at the 186 stations were next compared with the corresponding reported MMI levels. The resulting coefficient of correlation was found to be around 0.8. The point by point comparisons also showed very good agreement between equation III.1.2 and observed intensities. Equation (III.1.2) was then used to calculate the estimated MMI levels at those free-field sites in the new database for which no reported site intensities were available.

For scaling the Fourier amplitude spectra the following equation was employed:

 $\log_{10}[FS(T)] = b_1(T)\hat{I}_{MM} + b_2(T)h + b_3(T)v + b_4(T) , \qquad (III.1.3)$ which is the same as equation (III.1.1). Here \hat{I}_{MM} is the estimated MMI level at the site computed from equation (III.1.2) or the reported MMI level if available.

The distribution of the new database with respect to the levels of Modified Mercalli Intensity estimated or reported at the recording sites now becomes: MMI = II, 1 record, MMI = III, 8 records, MMI = IV, 15 records; MMI = V, 99 records; MMI = VI, 161 records; MMI = VII,

117 records; MMI = VIII, 30 records, MMI = IX, 2 records; and MMI = χ , 1 record.

The regression analysis was next performed on the new database with 1314 components of Fourier amplitudes FS(T), at 91 discrete periods T ranging from 0.04 to 15.0 sec. This procedure is identical to the procedure used with the old database (Trifunac, 1979; Trifunac and Lee, 1978). For completeness of this writing the details are repeated here briefly. As in the previous parts of the report, the data are first screened to minimize possible bias in the model. The data are partitioned into groups corresponding to MMI levels III, IV, V, VI, VII, VIII, IX and X. The data in each of these MMI levels are further subdivided according to the site classification parameter s = 0, 1 or 2. Depending on whether the recording component is horizontal or vertical, each of these subgroups is divided into 2 sets corresponding to horizontal (v = 0) and vertical (v = 1) components. The resulting data in each of the groups correspond to the Fourier spectral amplitudes from a specified MMI level for a specified site classification and with specified component orientation. The data points, $\log_{10} FS(T)$, within these groups are then arranged in increasing order according to their amplitudes. If the number of data points in a group is less than 19, all the data points are selected. If there are more than 19 points, at most 19 points are selected from among the ordered set of data so that they correspond uniformly, as close as possible, to the 5%, 10%,...,90% and 95% percentiles.

The resulting fitted coefficients at each period T resulting from linear regression have been denoted by $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$ and $\hat{b}_4(T)$, (equation (III.1.3) respectively.

III.2 THE REGRESSION COEFFICIENTS

Figure III.2.1 shows the smoothed coefficients $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$ and $\hat{b}_4(T)$ (solid lines) and the estimates of their 80%, 90% and 95% confidence intervals, represented by the corresponding dashed lines. Substituting these coefficients into equation (III.1.3) gives:

$$\log_{10} \hat{FS}(T) = \hat{b}_1(T) \hat{I}_{MM} + \hat{b}_2(T)h + \hat{b}_3(T)v + \hat{b}_4(T).$$
 (III.2.1)

 $\widehat{FS}(T)$ then represents the least squares estimate of the Fourier amplitude spectrum at period T for this model.

With FS(T) the corresponding Fourier amplitude spectrum computed from recorded accelerograms, the residuals, $\epsilon(T)$, were calculated as in Parts I and II of this report, where

$$\varepsilon(T) = \log_{10} FS(T) - \log_{10} \hat{FS}(T)$$
 (III.2.2)

As for Parts I and II, the assumption that $\epsilon(T)$ can be described by a normal distribution function with mean $\mu(T)$ and standard deviation $\sigma(T)$ is employed here. The probability $p(\epsilon,T)$ at period T that

$$\log_{10} FS(T) - \log_{10} \hat{FS}(T) \le \varepsilon(T)$$
 (III.2.3)

is then given by (equation (I.5.4) of Part I):

$$p(\varepsilon,T) = \frac{1}{\sigma(T)\sqrt{2\pi}} \int_{-\infty}^{\varepsilon(T)} \exp\left[-\frac{1}{2} \left(\frac{x-\mu(T)}{\sigma(T)}\right)^{2}\right] dx \qquad (III.2.4)$$

For a given residue, $\epsilon(T)$, at a particular period T, the actual probability $p^*(\epsilon,T)$ that $\epsilon(T)$ will not be exceeded can be evaluated by finding the fraction of residuals $\epsilon(T)$ (computed from the database at that particular period) which are smaller than a given value. Using (III.2.4), the estimated probability $\hat{p}(\epsilon,T)$ that $\epsilon(T)$ will not be

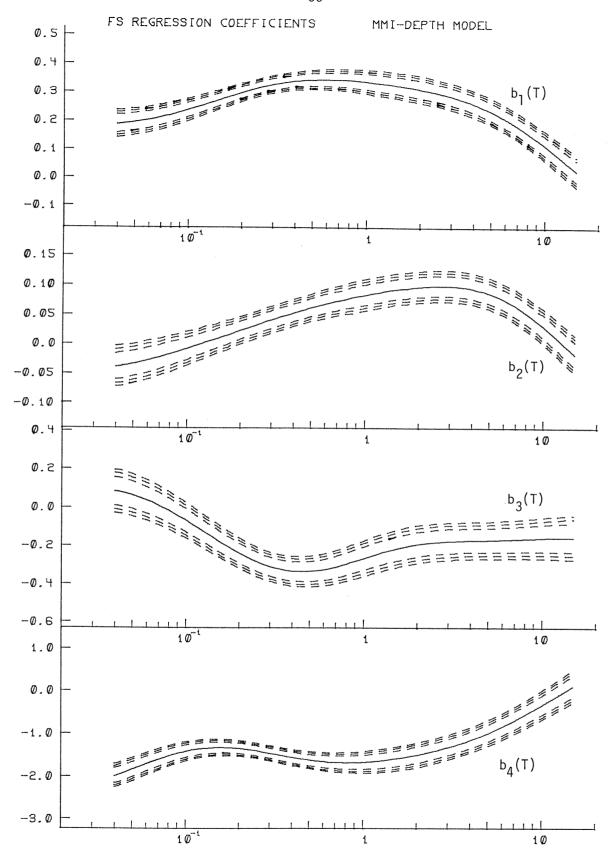


Figure III.2.1

exceeded can be compared with the above fractions. The Kolmogorov-Smirnov, KS(T), and the X^2 statistic, $X^2(T)$, can be computed to test the quality of fit of the normal distribution function in (III.2.4). A complete description of the steps involved in this and the formulae employed are given in Part I of this report and will not be repeated here.

Figure III.2.2 shows the plot of the amplitudes of the residuals corresponding to $p*(\epsilon,T)=0.1,\ 0.2,\dots,0.8$ and 0.9 for $\log_{10}FS(T)$. The nine sets of curves, plotted versus period T, from the bottom to the top of the plot correspond to the residual levels at each of the probability levels, 0.1 through 0.9. At each of the nine probability levels, the rough solid curve represents the actual calculated residuals at that particular level. The smoothed solid curve is obtained by smoothing the rough solid curve along the T-axis. The corresponding dashed curve is the estimated residual $\epsilon(T)$ at the particular probability level using equation (III.2.4).

It is of interest to compare this figure with the corresponding figure (Fig. I.5.3) in Part I of this report dealing with the scaling of FS(T) in terms of earthquake magnitude M and representative source to station distance Δ . There both the calculated and estimated residue levels of p* = 0.1 to p* = 0.9 range from ϵ = - 0.6 to ϵ = + 0.6 at the short period end to about ϵ = - 0.5 to ϵ = + 0.5 in the long period end (Part I, Fig. I.5.2). Correspondingly, here, they range from ϵ = - 0.7 to ϵ = + 0.7 at the short period end, and from ϵ = - 0.5 to ϵ = + 0.5 in the long period end. Since the smooth surface p*(ϵ ,T) from the nine smooth curves represents the spread of the observed data about the models given here (equation (III.1.3))

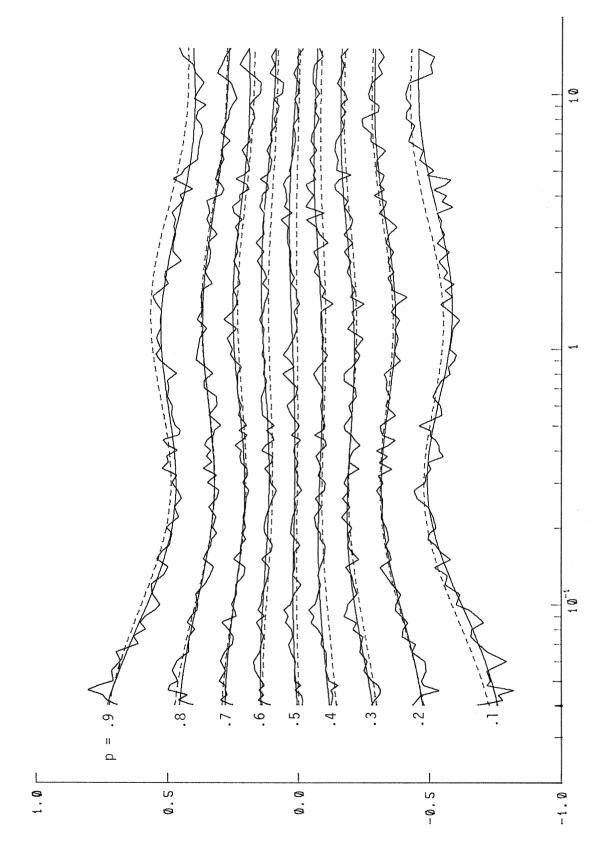


Figure III.2.2

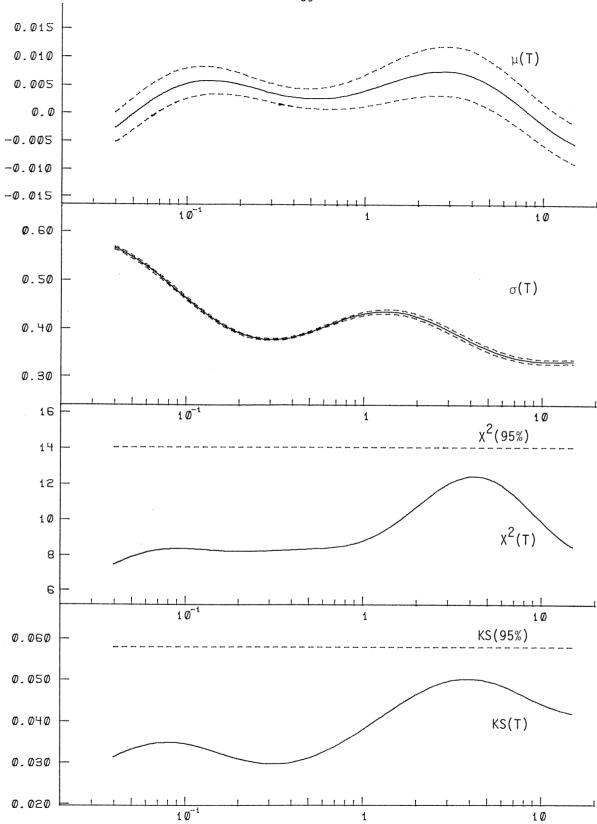


Figure III.2.3

TABLE III.2.1

$$\log_{10}FS(T) = b_1(T)I_{MM} + b_2(T)h + b_3(T)v + b_4(T)$$

and in Part I (equation I.5.1), this comparison shows that the uncertainties associated with the prediction of FS(T) in terms of MMI are very similar and certainly not much worse than those associated with the scaling of FS(T) in terms of M and Δ .

Figure III.2.3 shows a plot of the statistical parameters employed in the description of the residuals. The smooth amplitudes of $\hat{\mu}(T)$ and $\hat{\sigma}(T)$ of equation (III.2.4) and their 95% confidence intervals are shown in the top 2 plots of the figure, respectively. The two full curves in the bottom of the figure show the smoothed amplitudes of the computed X^2 , $X^2(T)$ and Komolgorov-Smirnov, KS(T), statistics, respectively. The dashed lines are their corresponding 95% cutoff levels. It is seen that in the whole period range considered (0.04 sec to 15 sec), both the X^2 and K-S tests fail to reject the hypothesis that the distribution is normal.

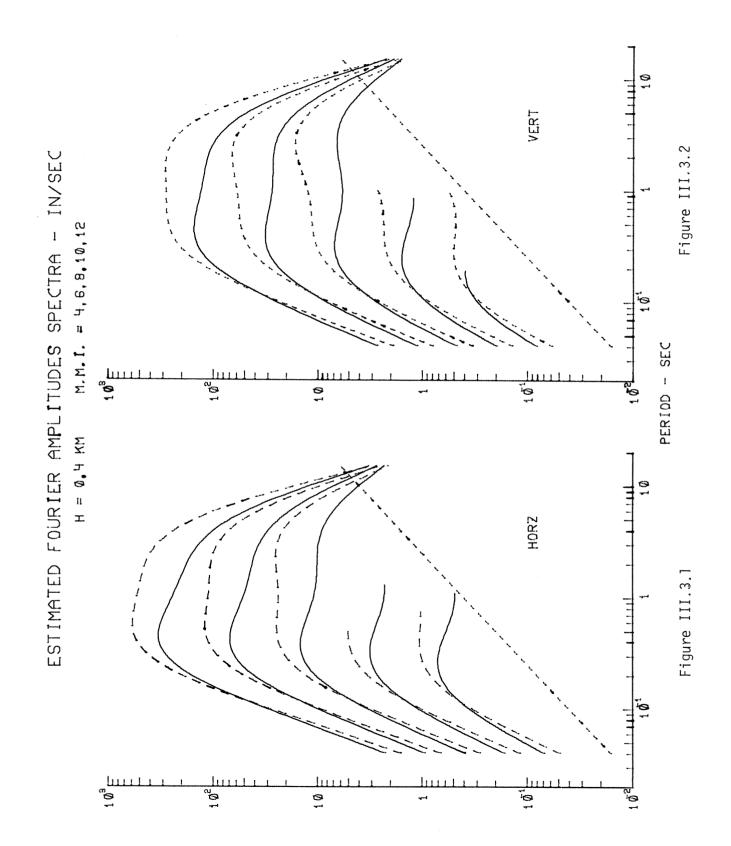
Table III.2.1 gives, for 12 periods between T = 0.04 sec and T = 14 sec, the amplitudes of the smoothed regression coefficients: $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$, $\hat{b}_4(T)$, the nine smoothed calculated residue levels corresponding to p*(ϵ ,T) = 0.1 through 0.9, the smoothed coefficients $\hat{\mu}(T)$, $\hat{\sigma}(T)$ of the normal distribution function in (III.2.4), and finally, the X² and the Komolgorov-Smirnov statistics.

III.3 THE ESTIMATED FOURIER SPECTRA

Figures III.3.1 and III.3.2 present examples of the Fourier amplitude spectra, FS(T), computed from equation (III.1.3) for $p(\varepsilon,T)=0.5$, for MMI levels IV, VI, VIII, X and XII. Figure III.3.1 is for horizontal motion (v = 0) while Figure III.3.2 is for vertical motion (v = 1). The solid lines in both figures correspond to the depth of sediments h = 0, while the dashed lines correspond to h = 4 km. The diagonal dashed lines at the bottom of each graph represent the average Fourier amplitudes of the digitization noise. The plot of each spectrum is presented only for those periods where the low signal-to-noise ratio in the data does not distort the estimates of small spectral amplitudes.

The overall trends of the computed FS(T) spectral amplitudes in these figures are in many ways similar to those of the same model in our previous analysis (Trifunac and Lee, 1978). The results can be considered representative of the observed shaking for MMI levels up to about VIII. The curves plotted for MMI = X and XII are presented here only for completeness, and represent an extrapolation based on the currently available data.

It has been noted in the previous analysis of the same model (Trifunac, 1979; Trifunac and Lee, 1978), that for the same MMI, spectral amplitudes tend to be slightly higher for h=0 km than for say h=4 km in the short period range (up to 0.2 sec), with this trend being reversed for long periods. Note that the same trend is observed here in Figures III.3.1 and III.3.2, but to a smaller extent. Here the amplitudes are slightly higher for h=0 km than for h=4 km, only in



AC041 PAC0IMA DAM, 1971 COMP HORZ MMI = 10. DEPTH = 0.FT V = 0.

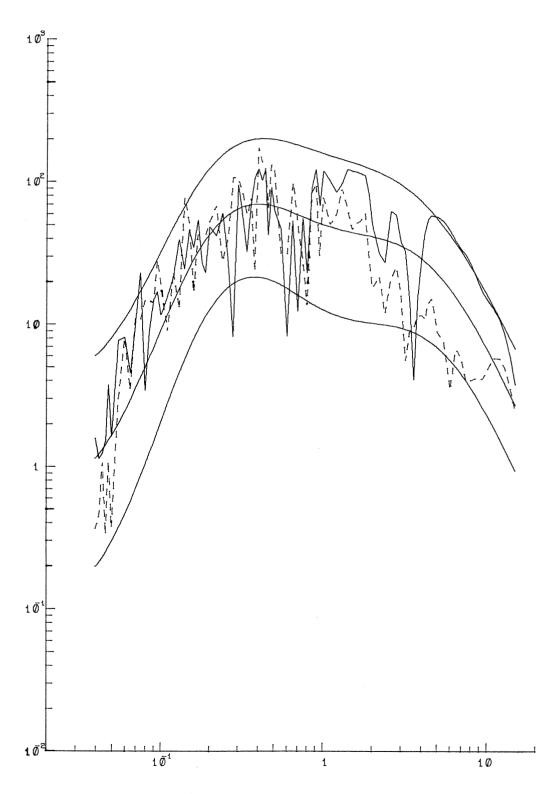


Figure III.3.3

AC041 PAC0IMA DAM, 1971 COMP DOWN MMI = 10. DEPTH = 0.FT V = 1.

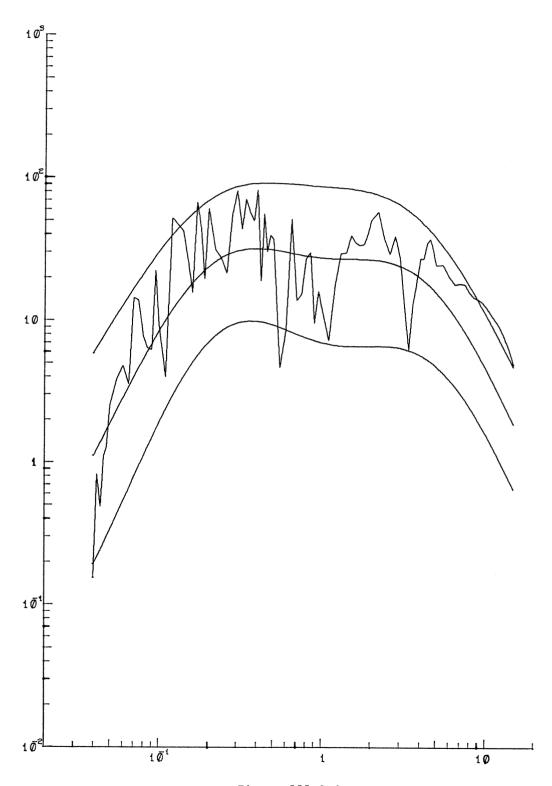


Figure III.3.4

AA001 EL CENTRO, 1940 COMP HORZ MMI = 8. DEPTH =15000.FT V = 0.

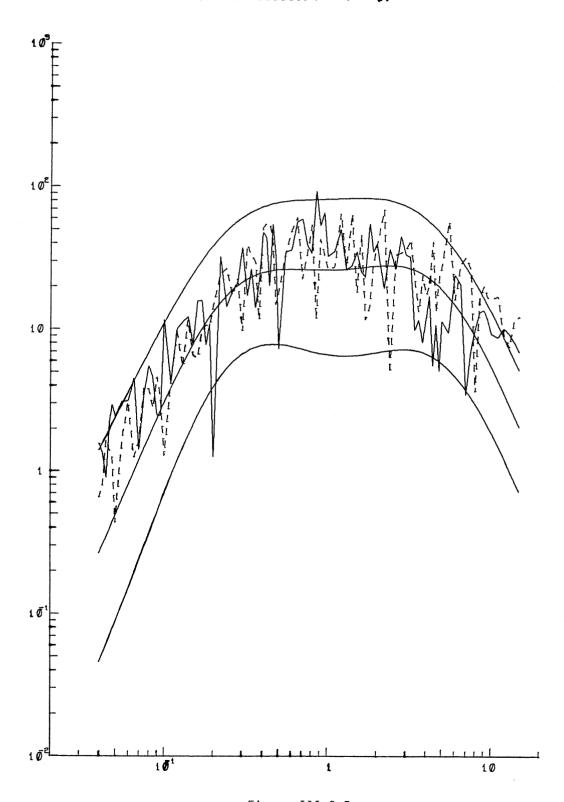


Figure III.3.5

AA001 EL CENTRO, 1940 COMP VERT MMI = 8. DEPTH =15000.FT V = 1.

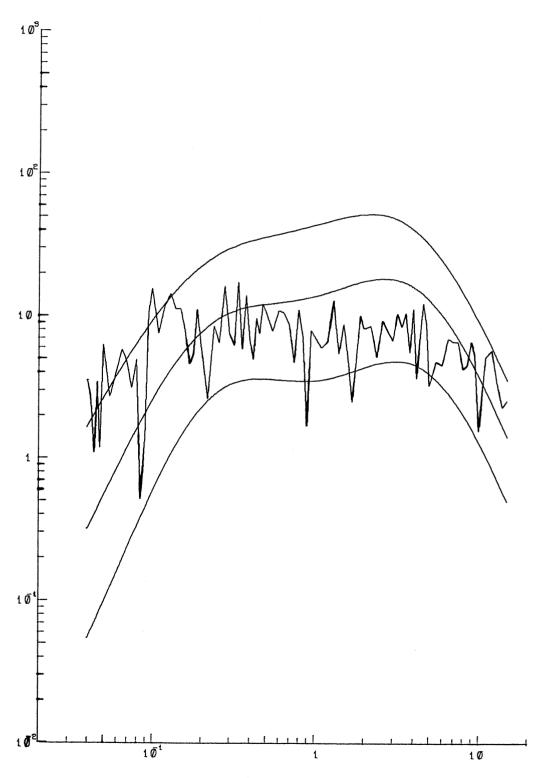


Figure III.3.6

the short period range of up to about 0.1 sec, beyond which the trend is reversed. It is also of interest to compare these figures with the corresponding figures (Figures I.6.1 and I.6.2) in Part I of this report dealing with the scaling of FS(T) in terms of M and Δ . It has been found there that the effect of the depth of sediments, h, beneath the recording station is important only from intermediate to long periods and is negligible at short periods. There, for the short periods, no trend reversal has been observed.

A comparison of the corresponding horizontal and vertical amplitudes in Figures III.3.1 and III.3.2 also shows that the vertical amplitudes are smaller than the horizontal amplitudes, except at short periods, a trend which is consistent with all previous analyses.

Figures III.3.3 and III.3.4 compare the amplitudes of Fourier spectra computed from accelerations recorded at Pacoima Dam during the San Fernando earthquake of 1971, with spectral amplitudes computed from equation (III.1.3) for p = 0.1, 0.5 and 0.9 with MMI = 10, h = 0 ft. and for the vertical and two horizontal components. The agreement between the observed and estimated amplitudes is good for all three components. Figures III.3.5 and III.3.6 compare the FS amplitudes computed from accelerations recorded during the Imperial Valley, California earthquake of 1940, in El Centro, with the estimated amplitudes using MMI = 8 and h = 15000 ft for both the vertical and horizontal components. The agreement is good for the two horizontal components and fair for the vertical component.

This completes the description of the preliminary model for scaling FS(T) in terms of MMI, h and v.

PART IV: <u>SCALING OF FOURIER SPECTRA IN TERMS OF MMI, s and v</u> IV.1 THE SCALING RELATION

Part IV of this report continues the description of the preliminary empirical model for scaling Fourier amplitude spectra of strong ground motion in terms of Modified Mercalli Intensity (MMI) at the site and local geology. As in Part II of this report, this part of the analysis replaces the depth of sedimentary deposits h, employed as site characterization in the previous Part III, by the corresponding site parameter s = 0, 1 and 2. After modifying the scaling relation of Part III, equation (III.1.3), the scaling relation now takes the form

$$\log_{10} FS(T) = b_1(T) \hat{I}_{MM} + b_2(T) s + b_3(T) v + b_4(T) , \qquad (IV.1.1)$$
 with all parameters defined as before. $b_2(T)$ is now the coefficient associated with the site parameter s.

The scaling functions $b_1(T)$ through $b_4(T)$ are determined again through a regression analysis of the new database of 1314 components of spectral amplitudes, FS(T), at 91 discrete periods T ranging from 0.04 sec to 15.0 sec. As in the previous Part III of this report, the data are first screened for possible bias in the model. All procedures in data preparation and selection, and the steps of regression analysis employed here are identical to those in Part III of this report, and so their description will not be repeated here.

The coefficients at each period T resulting from linear regression have been denoted by $\hat{b}_1(T)$, $\hat{b}_2(T)$, $\hat{b}_3(T)$ and $\hat{b}_4(T)$ (equation (IV.1.1)), respectively. Much of the format of the description in the sections to follow will almost be identical to that in Part III of this work. The reader may refer to the corresponding sections of Part III for a more detailed description.

IV.2 THE REGRESSION COEFFICIENTS

Figure IV.2.1 shows the smoothed coefficients \hat{b}_1 (T) through \hat{b}_4 (T) (solid lines) together with the estimates of their 80%, 90% and 95% confidence intervals (dashed lines). Comparison of this figure with the corresponding Figure III.2.1 in the previous Part III of this report shows that the functions \hat{b}_1 (T), \hat{b}_3 (T) and \hat{b}_4 (T) as given respectively by the top and bottom two graphs are almost identical. These functions correspond to the same respective parameters, I_{MM} , v and 1 in the scaling relations and their similarity again demonstrates the stability of the two regression models used for scaling. The functions \hat{b}_2 (T) as given by the second graph from the top in both figures are opposite in sign. This again is consistent for the two models since s=2 corresponds to s=2 corresponds to s=3 km (basement rock), while s=3 corresponds to s=3 km (alluvium).

Figure IV.2.2 shows the plot of the residual levels corresponding to $p*(\epsilon,T) = 0.1, 0.2,...,0.8, 0.9$ for $log_{10}FS(T)$. Refer to the same Figure III.2.2 in Part III of this report for a complete description of each set of these curves.

It is again of interest to compare the two figures IV.2.2 and III.2.2. Since the two smooth surfaces $p*(\epsilon,T)$ represent the spread of the observed data about their corresponding models, which differ only in the characterization of local site geology, the resemblance of the two figures again shows that the uncertainties associated with the characterization of local geology in terms of site conditions s=0, 1 and 2 are not much greater than those associated with the site characterization in terms of depth of sedimentary deposits.

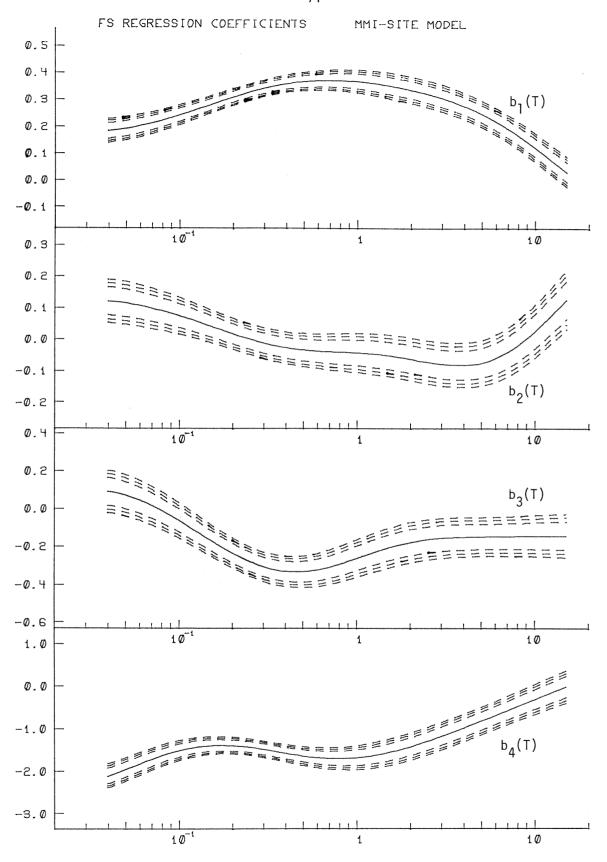


Figure IV.2.1

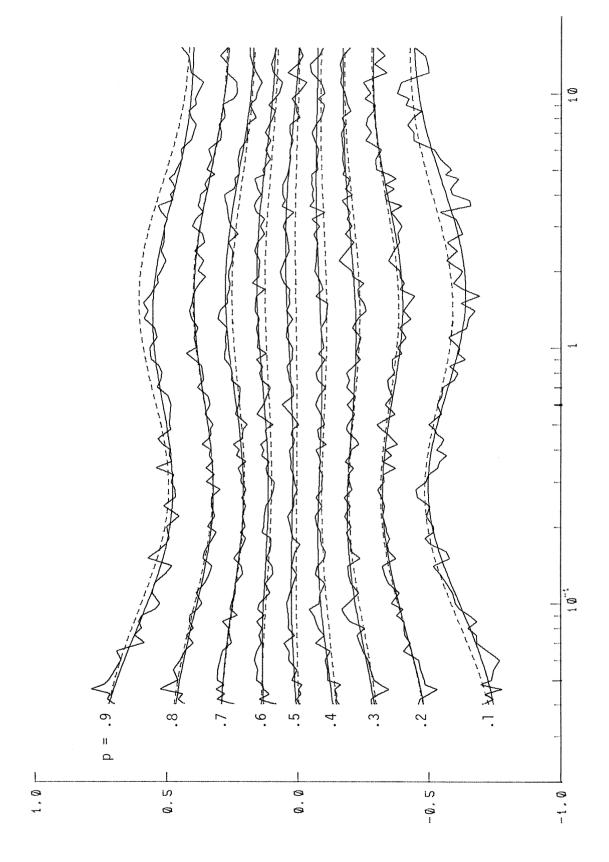


Figure IV.2.2

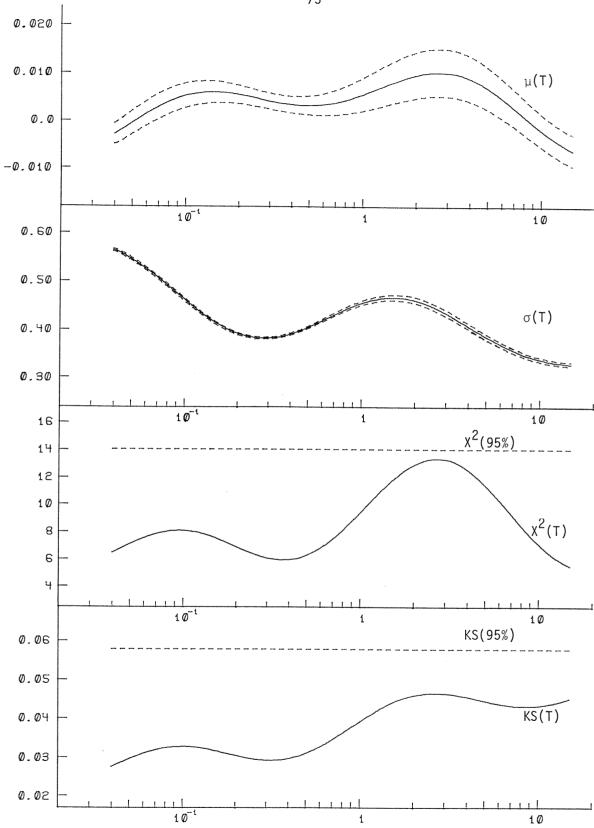


Figure IV.2.3

TABLE IV.2.1

	14.00	.045 .109 143	445 287 172 003 .088 .169 .266	006 .329 5.678
$\log_{10} FS(T) = b_1(T)I_{MM} + b_2(T)s + b_3(T)v + b_4(T)$	7.50	. 179 - 029 - 144 - 519	484 297 174 071 . 020 . 108 . 194 . 298	.002 .353 8.821
	4.40	. 263 081 148 911	551 333 184 068 .039 .134 .342	.008 .397 12.090
	2.80	.310 079 157	609 372 199 072 .048 .151 .262	.010 .439 13.381
	1.60	.347 056 200 -1.542	638 399 082 .044 .156 .277 .397	.009 .467 12.095
	06•	.369 044 277 -1.709	502 384 220 090 031 142 258	.005 .450 8.829
	• 50	.368 036 334 -1.659	530 345 205 089 .021 .122 .224 .343	.003 .405 6.306
	.34	.352 021 323 -1.540	500 326 194 083 019 113 209	.004 .385 5.962
	.19	.306 .019 227 -1.397	518 332 190 077 116 213	.006 .394 7.008
	.11	.253 .065 090 -1.476	599 375 084 .024 .128 .375	.006 .448 8.047
	.065	.209 .101 .028 -1.774	689 433 107 .019 .273 .428	.002 .516 7.750
	.040	.184 .121 .091 -2.118	744 478 291 133 .008 .131 .290 .465	ICS: - 003 - 565 5.504
	PERIOD, T (SEC)	COEFFICIENTS: $b_1(T)$ $b_2(T)$ $b_3(T)$ $b_4(T)$	RESIDUES: p = .1 p = .2 p = .4 p = .5 p = .6 p = .7 p = .8	RESIDUE STATISTICS: µ(T) α(T) X ² (T) KS(T)

Figure IV.2.3 shows the plot of the statistical parameters employed in the description of the residuals, namely, $\hat{\mu}(T)$, $\hat{\sigma}(T)$, $\chi^2(T)$ and KS(T), from top to bottom. Comparison with the Figure III.2.3 in Part III again shows the degree of the resemblance of this model and that discussed in Part III.

Table IV.2.1 gives, for 12 periods between T = 0.04 sec and T = 14 sec, the four coefficients, $\hat{b}_1(T)$ through $\hat{b}_4(T)$, the nine residue levels corresponding to $p*(\varepsilon,T)=0.1$ through 0.9, the coefficients $\hat{\mu}(T)$ and $\hat{\sigma}(T)$ of the normal distribution and finally the $X^2(T)$ and KS(T) statistics.

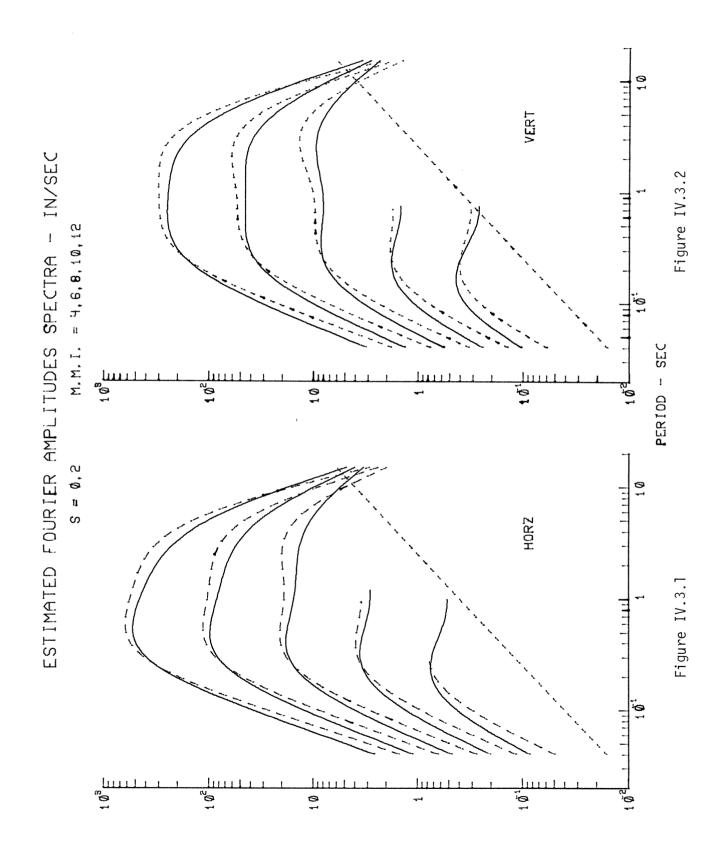
IV.3 THE ESTIMATED FOURIER SPECTRA

Figures IV.3.1 and IV.3.2 present examples of the Fourier amplitude spectra, FS(T), computed from equation (IV.1.1) for $p(\epsilon,T)=0.5$, for MMI levels IV, VI, VIII, X and XII. Figure IV.3.1 is for horizontal motion (v = 0) while Figure IV.3.2 is for vertical motion (v = 1). The solid lines in both figures correspond to the site condition s = 2 while the dashed lines correspond to s = 0. The diagonal dashed lines again represent the empirical average Fourier amplitudes of digitization noise.

Comparison of these figures with the corresponding Figures III.3.1 and III.3.2 of Part III of this report again shows great similarity, and hence the similar conclusions can be drawn from these figures.

Figures IV.3.3 and IV.3.4 compare the amplitudes computed from the 1971 Pacoima Dam acceleration with those calculated for p=0.1, 0.5 and 0.9, MMI = X and s=2. Similarly, Figures IV.3.5 and IV.3.6 compare those computed from the 1940 El Centro acceleration with those calculated for MMI = VIII and s=0. The agreement ranges from fair to good.

This completes the description of the Part IV of the scaling of FS(T) in terms of MMI, s and v.



AC041 PACOIMA DAM 1971 COMP S16E S74W MMI = 10. SITE = 2. V = 0.

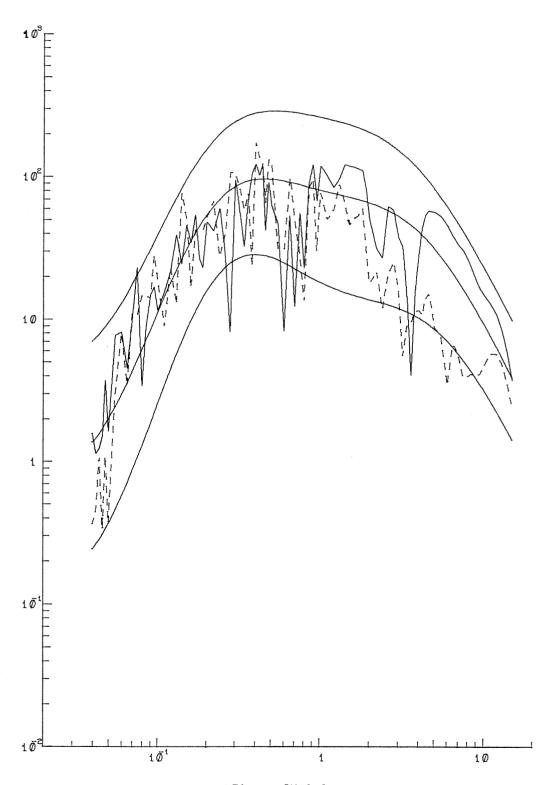


Figure IV.3.3

ACØ41 PACOIMA DAM, 1971 COMP DOWN MMI = 10. SITE = 2. V = 1.

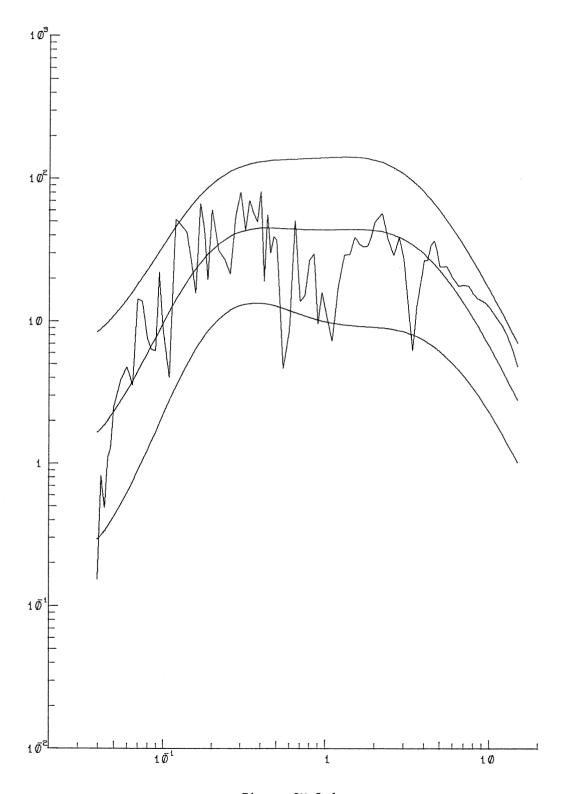


Figure IV.3.4

AA001 EL CENTRO, 1940 COMP NS, EW MMI = 8. SITE = 0. V = 0.

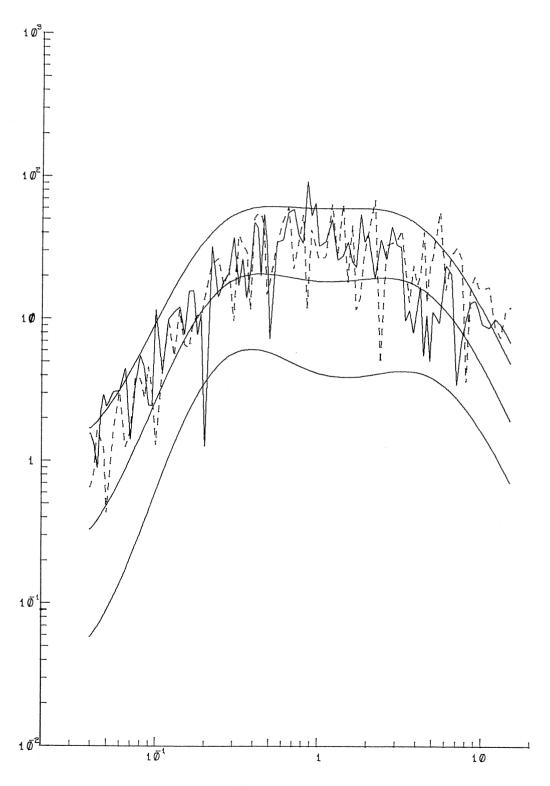


Figure IV.3.5

AA001 EL CENTRO, 1940 COMP VERT MMI = 8. SITE = 0. V = 1.

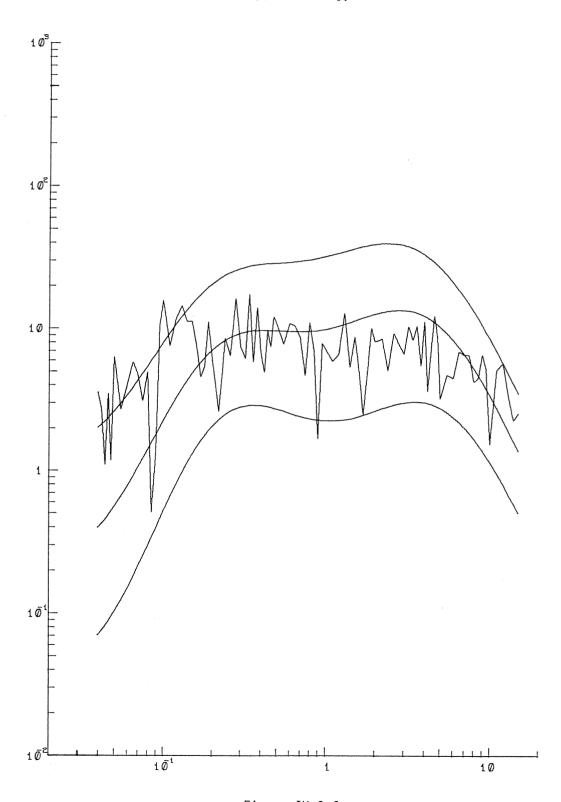


Figure IV.3.6

CONCLUSIONS

The results of this study can be summarized as follows:

- 1. On the logarithmic scale the Fourier spectrum amplitudes grow linearly for small magnitudes (M \leq 3). This growth rate reduces for intermediate magnitudes and stops for M \gtrsim 8, in agreement with our earlier studies (Trifunac, 1976b).
- 2. For intermediate and long periods Fourier spectrum amplitudes recorded on alluvium are larger than those recorded on basement rock. At high frequencies, in contrast to our previous work (Trifunac, 1976b), $b_2(T)$ in Model I remains positive though not significantly different from zero. While this may be a consequence of the overall model differences and the fluctuations of the "constant" term $b_5(T)$, this suggests that our earlier discussions based on d(T) in Trifunac (1976b) being negative may have to be modified. In most of the earlier analyses we found $b_2(T)$, or its equivalent, to be negative for frequencies higher than about 5 Hz, but with small amplitudes which are not significantly different from zero. The overall shape of $b_2(T)$ function or its equivalent, however, has been very stable in present as well as in all earlier analyses.
- 3. The scaling function $b_3(T)$ in all four models which reflects the differences between horizontal and vertical spectral amplitudes is in excellent agreement with e(T) in Trifunac (1976b). Again for high frequencies the vertical spectral amplitudes tend to be equal to slightly larger than the horizontal amplitudes.
- 4. The distribution of Fourier spectrum amplitudes about the estimated model can be described adequately by the log-normal distribution function. Except for periods longer than about 10 sec, the quality of

this fit is excellent. For periods longer than 0.2 sec the standard deviation of this distribution function is about 0.35.

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